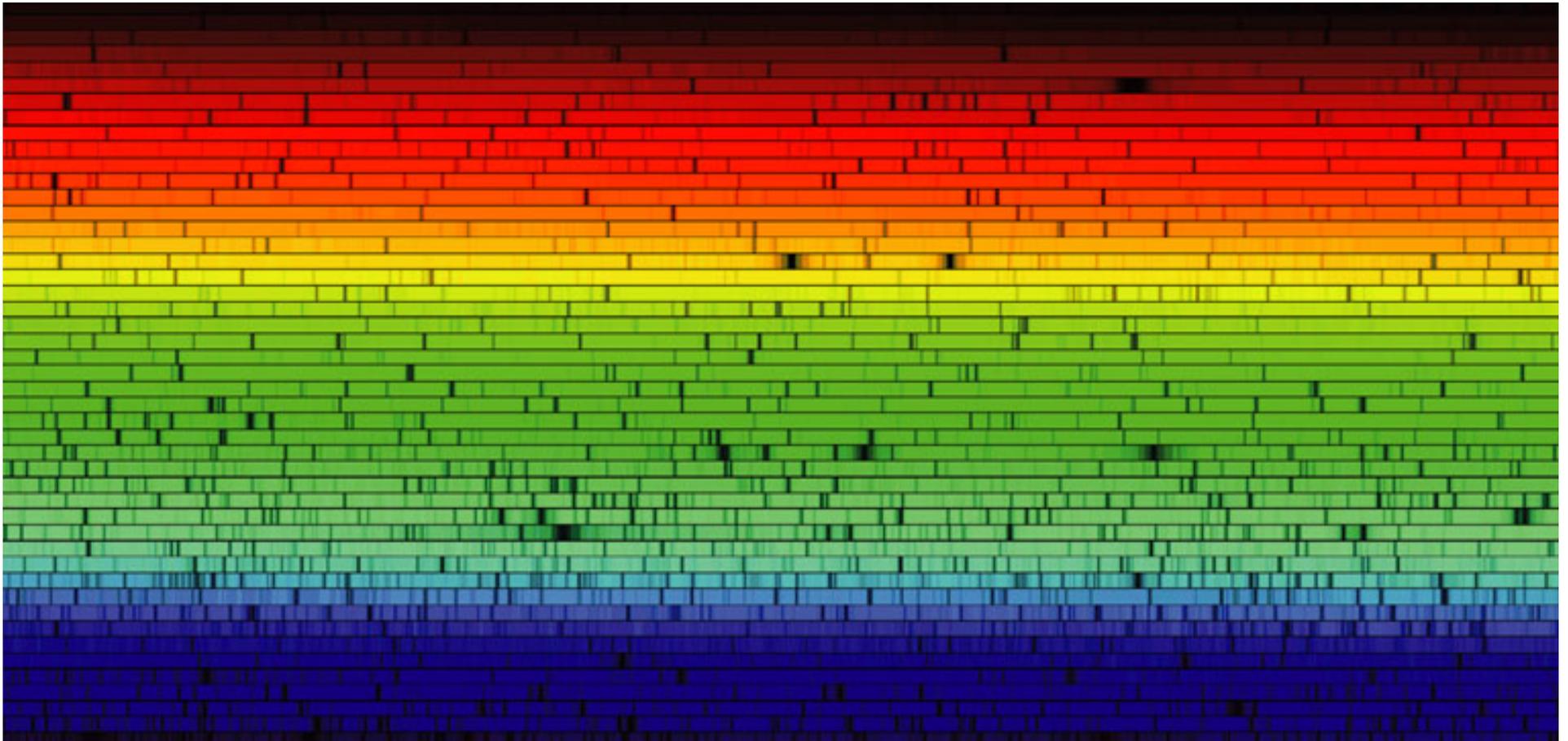
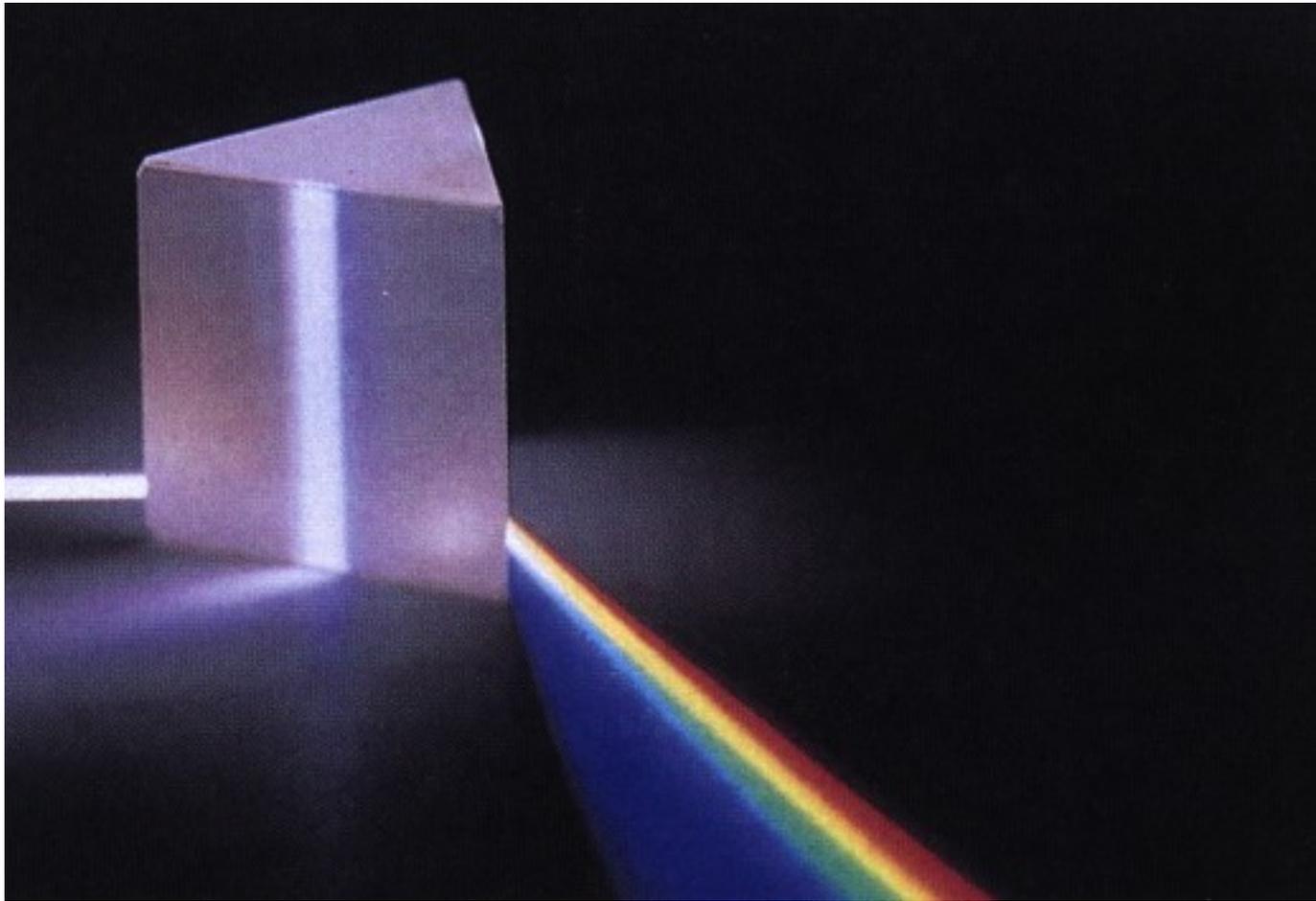


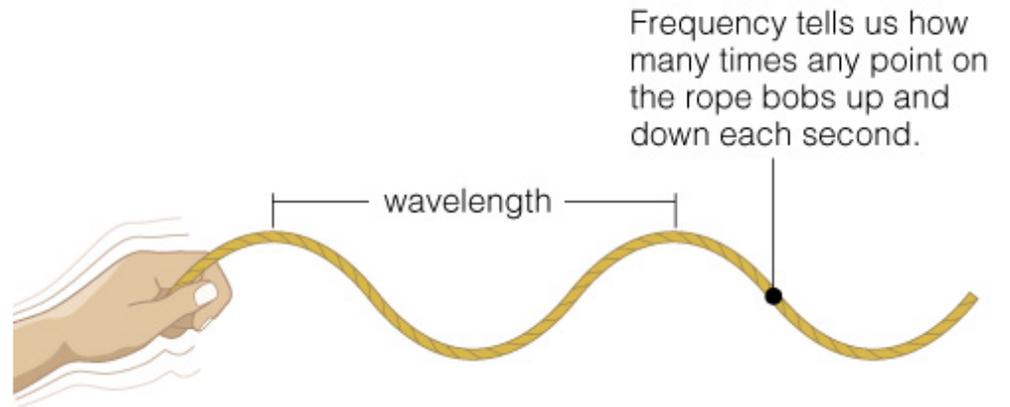
# Radiation, Matter and Energy



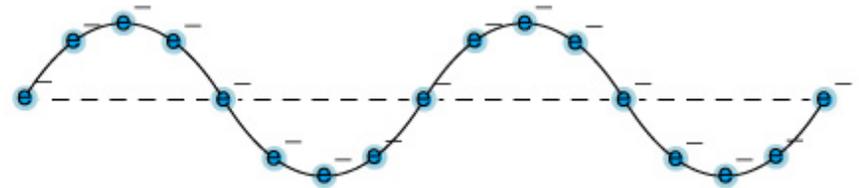
# What is light?



# Light is an electromagnetic wave



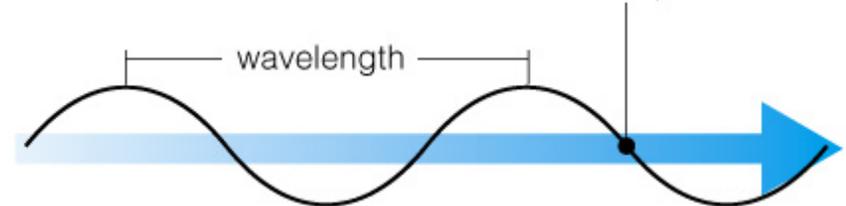
**a**



**b**

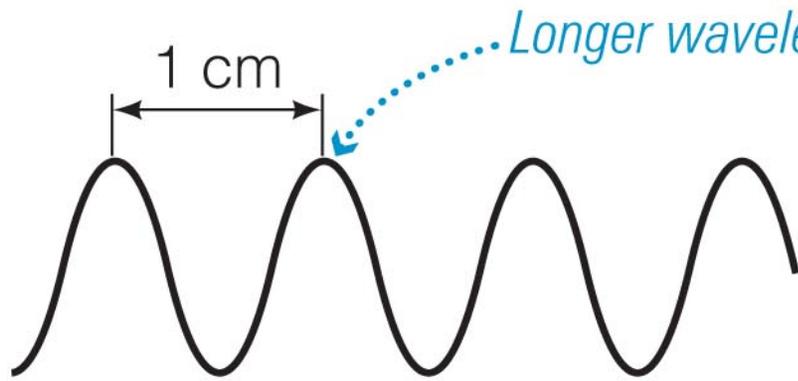
Wavelength is the distance between adjacent peaks of the electric field.

Frequency is the number of times each second that the electric field peaks at any point.



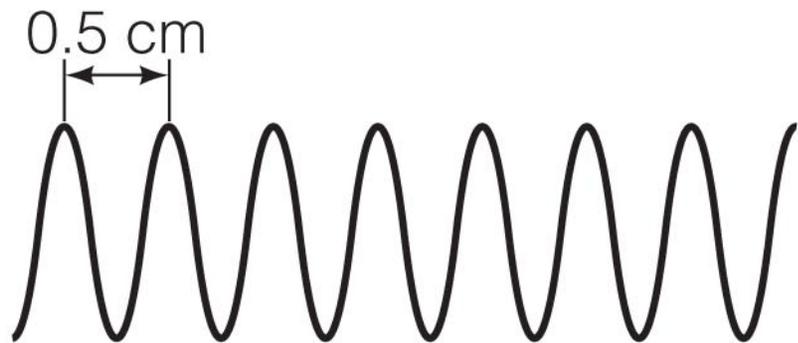
All light travels with speed  $c = 300,000$  km/s.

**c**

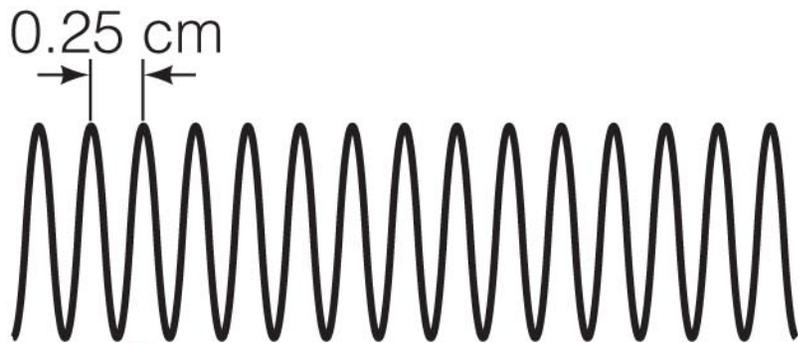


wavelength = 1 cm,  
frequency = 30 Ghz

*Longer wavelength means lower frequency.*



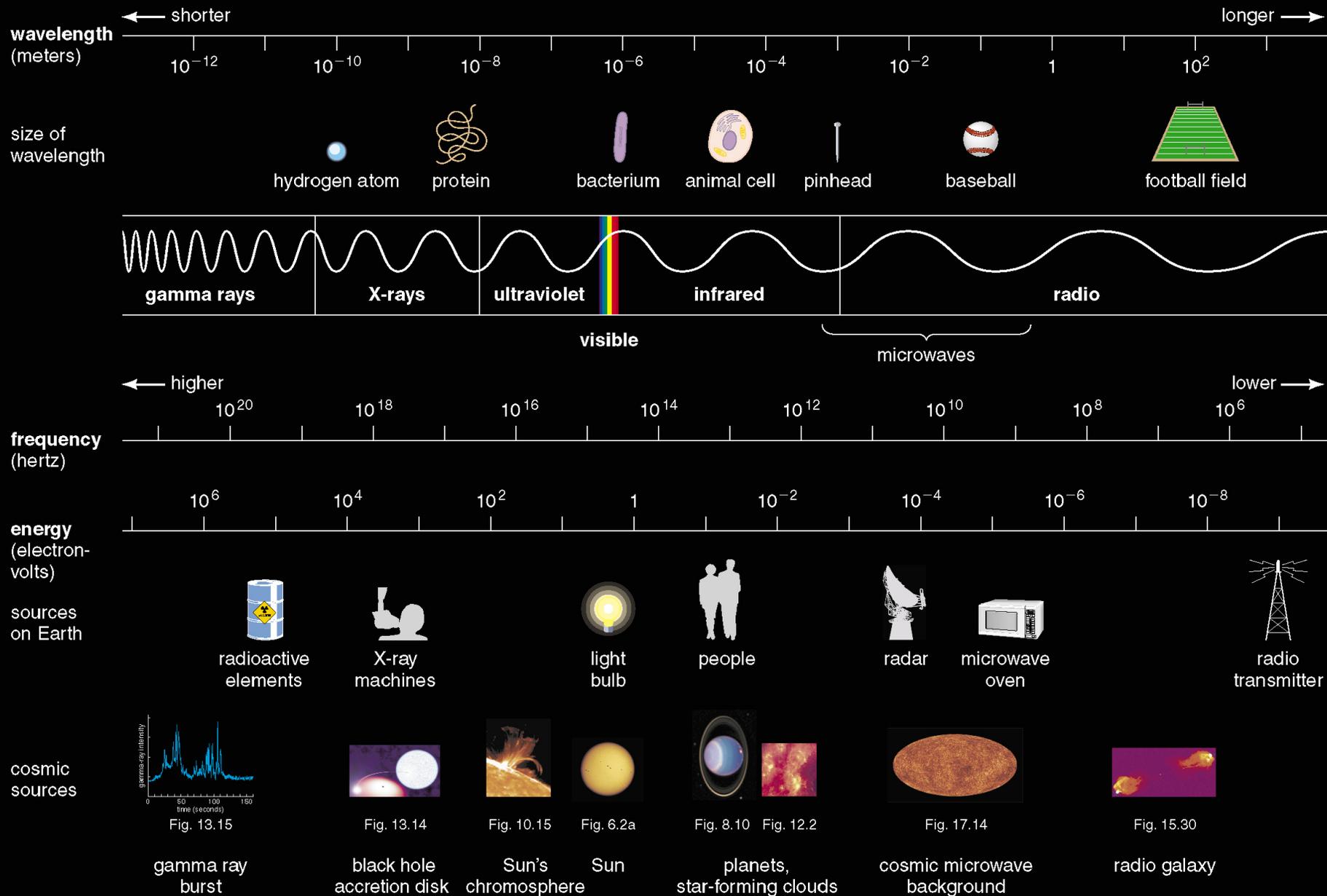
wavelength =  $\frac{1}{2}$  cm,  
frequency =  $2 \times 30$  Ghz = 60 Ghz



wavelength =  $\frac{1}{4}$  cm,  
frequency =  $4 \times 30$  Ghz = 120 Ghz

*Shorter wavelength means higher frequency.*

# The Electromagnetic Spectrum



# Light is also a particle

**Photons:** “pieces” of light, each with precise wavelength, frequency, and energy. Think of photons as tiny bullets

Energy is proportional to frequency

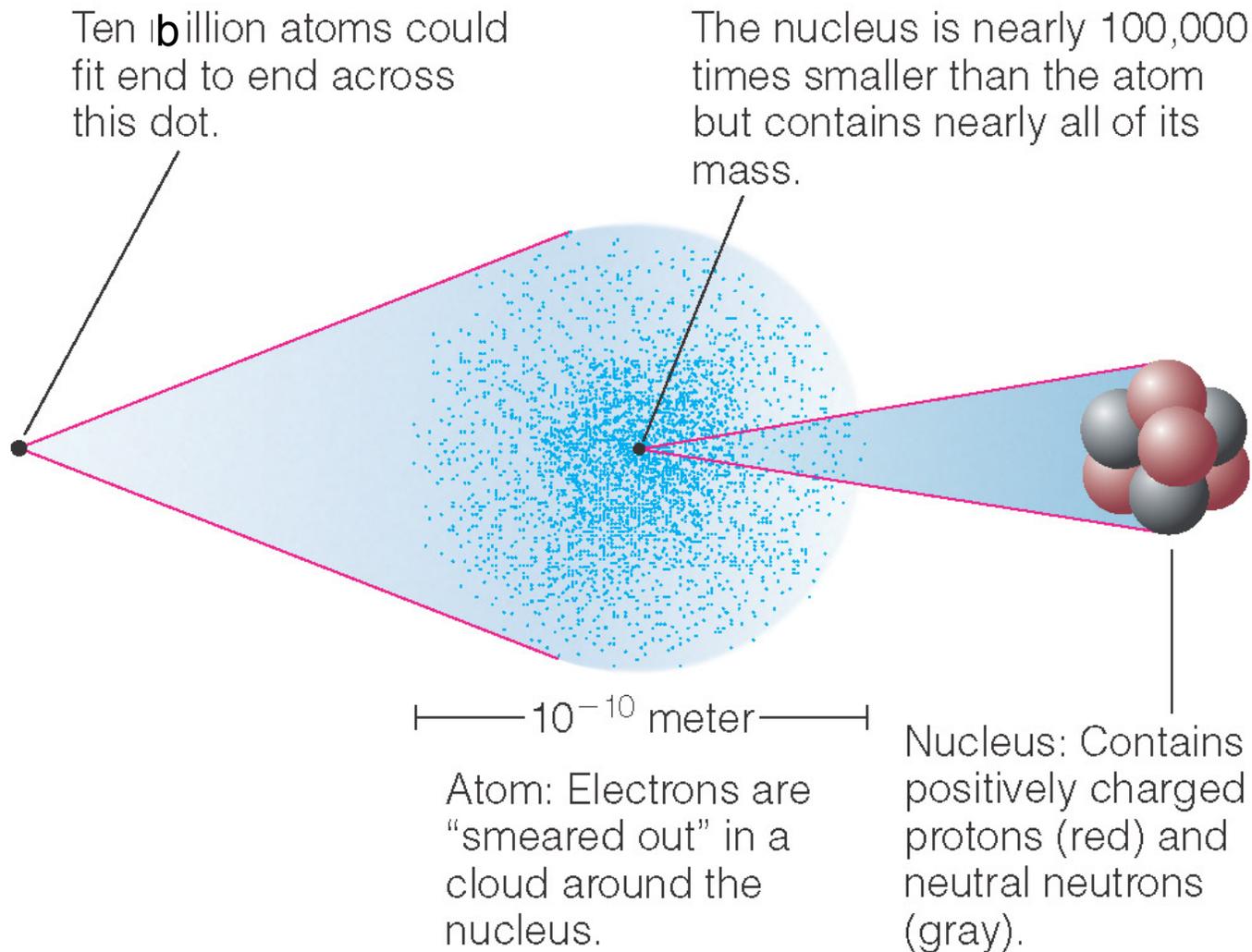
Within the visible spectrum, **blue light** has higher energy than **red light**

Within the electromagnetic spectrum, X-rays have the highest energy, followed by UV, visible light, IR, and radio

***Remember:** Light is just one form of electromagnetic wave of energy, the kind we can detect with our eyes.*

# What is matter?

## Atomic structure:

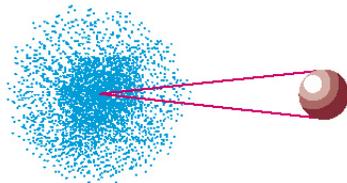


# Atomic Terminology

- **Atomic Number** = # of protons in nucleus
- **Atomic Mass Number** = # of protons + neutrons

atomic number = number of protons  
atomic mass number = number of protons + neutrons

**Hydrogen ( ${}^1\text{H}$ )**

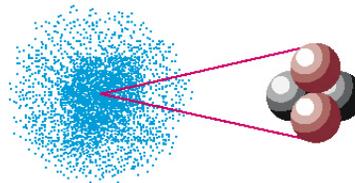


atomic number = 1

atomic mass number = 1

(1 electron)

**Helium ( ${}^4\text{He}$ )**

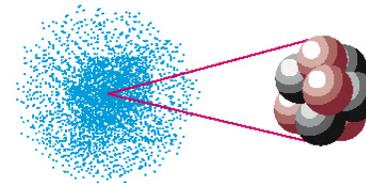


atomic number = 2

atomic mass number = 4

(2 electrons)

**Carbon ( ${}^{12}\text{C}$ )**



atomic number = 6

atomic mass number = 12

(6 electrons)

The number of electrons in a neutral atom equals its atomic number.

# Atoms organized on Periodic Table

- Atomic Number = # of protons in nucleus

hydrogen 1 <b>H</b> 1.0079																						helium 2 <b>He</b> 4.0026	
lithium 3 <b>Li</b> 6.941	beryllium 4 <b>Be</b> 9.0122											boron 5 <b>B</b> 10.811	carbon 6 <b>C</b> 12.011	nitrogen 7 <b>N</b> 14.007	oxygen 8 <b>O</b> 15.999	fluorine 9 <b>F</b> 18.998	neon 10 <b>Ne</b> 20.180						
sodium 11 <b>Na</b> 22.990	magnesium 12 <b>Mg</b> 24.305											aluminium 13 <b>Al</b> 26.982	silicon 14 <b>Si</b> 28.086	phosphorus 15 <b>P</b> 30.974	sulfur 16 <b>S</b> 32.065	chlorine 17 <b>Cl</b> 35.453	argon 18 <b>Ar</b> 39.948						
potassium 19 <b>K</b> 39.098	calcium 20 <b>Ca</b> 40.078	scandium 21 <b>Sc</b> 44.956	titanium 22 <b>Ti</b> 47.867	vanadium 23 <b>V</b> 50.942	chromium 24 <b>Cr</b> 51.996	manganese 25 <b>Mn</b> 54.938	iron 26 <b>Fe</b> 55.845	cobalt 27 <b>Co</b> 58.933	nickel 28 <b>Ni</b> 58.693	copper 29 <b>Cu</b> 63.546	zinc 30 <b>Zn</b> 65.39	gallium 31 <b>Ga</b> 69.723	germanium 32 <b>Ge</b> 72.61	arsenic 33 <b>As</b> 74.922	selenium 34 <b>Se</b> 78.96	bromine 35 <b>Br</b> 79.904	krypton 36 <b>Kr</b> 83.80						
rubidium 37 <b>Rb</b> 85.468	strontium 38 <b>Sr</b> 87.62	yttrium 39 <b>Y</b> 88.906	zirconium 40 <b>Zr</b> 91.224	niobium 41 <b>Nb</b> 92.906	molybdenum 42 <b>Mo</b> 95.94	technetium 43 <b>Tc</b> [98]	ruthenium 44 <b>Ru</b> 101.07	rhodium 45 <b>Rh</b> 102.91	palladium 46 <b>Pd</b> 106.42	silver 47 <b>Ag</b> 107.87	cadmium 48 <b>Cd</b> 112.41	indium 49 <b>In</b> 114.82	tin 50 <b>Sn</b> 118.71	antimony 51 <b>Sb</b> 121.76	tellurium 52 <b>Te</b> 127.60	iodine 53 <b>I</b> 126.90	xenon 54 <b>Xe</b> 131.29						
caesium 55 <b>Cs</b> 132.91	barium 56 <b>Ba</b> 137.33	57-70 *	lutetium 71 <b>Lu</b> 174.97	hafnium 72 <b>Hf</b> 178.49	tantalum 73 <b>Ta</b> 180.95	tungsten 74 <b>W</b> 183.84	rhenium 75 <b>Re</b> 186.21	osmium 76 <b>Os</b> 190.23	iridium 77 <b>Ir</b> 192.22	platinum 78 <b>Pt</b> 195.08	gold 79 <b>Au</b> 196.97	mercury 80 <b>Hg</b> 200.59	thallium 81 <b>Tl</b> 204.38	lead 82 <b>Pb</b> 207.2	bismuth 83 <b>Bi</b> 208.98	polonium 84 <b>Po</b> [209]	astatine 85 <b>At</b> [210]	radon 86 <b>Rn</b> [222]					
francium 87 <b>Fr</b> [223]	radium 88 <b>Ra</b> [226]	89-102 * *	lawrencium 103 <b>Lr</b> [262]	rutherfordium 104 <b>Rf</b> [261]	dubnium 105 <b>Db</b> [262]	seaborgium 106 <b>Sg</b> [266]	bohrium 107 <b>Bh</b> [264]	hassium 108 <b>Hs</b> [269]	meitnerium 109 <b>Mt</b> [268]	ununnilium 110 <b>Uun</b> [271]	unununium 111 <b>Uuu</b> [272]	ununbium 112 <b>Uub</b> [277]		ununquadium 114 <b>Uuq</b> [289]									

\* Lanthanide series

lanthanum 57 <b>La</b> 138.91	cerium 58 <b>Ce</b> 140.12	praseodymium 59 <b>Pr</b> 140.91	neodymium 60 <b>Nd</b> 144.24	promethium 61 <b>Pm</b> [145]	samarium 62 <b>Sm</b> 150.36	europium 63 <b>Eu</b> 151.96	gadolinium 64 <b>Gd</b> 157.25	terbium 65 <b>Tb</b> 158.93	dysprosium 66 <b>Dy</b> 162.50	holmium 67 <b>Ho</b> 164.93	erbium 68 <b>Er</b> 167.26	thulium 69 <b>Tm</b> 168.93	ytterbium 70 <b>Yb</b> 173.04
actinium 89 <b>Ac</b> [227]	thorium 90 <b>Th</b> 232.04	protactinium 91 <b>Pa</b> 231.04	uranium 92 <b>U</b> 238.03	neptunium 93 <b>Np</b> [237]	plutonium 94 <b>Pu</b> [244]	americium 95 <b>Am</b> [243]	curium 96 <b>Cm</b> [247]	berkelium 97 <b>Bk</b> [247]	californium 98 <b>Cf</b> [251]	einsteinium 99 <b>Es</b> [252]	fermium 100 <b>Fm</b> [257]	mendelevium 101 <b>Md</b> [258]	nobelium 102 <b>No</b> [259]

\*\* Actinide series

# Atomic Terminology

- **Isotope:** same # of protons but different # of neutrons. ( $^4\text{He}$ ,  $^3\text{He}$ )

Different isotopes of a given element contain the same number of protons but different numbers of neutrons.

## Isotopes of Carbon

carbon-12



$^{12}\text{C}$

(6 protons  
+ 6 neutrons)

carbon-13



$^{13}\text{C}$

(6 protons  
+ 7 neutrons)

carbon-14

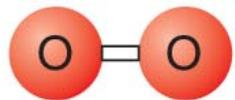


$^{14}\text{C}$

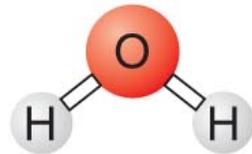
(6 protons  
+ 8 neutrons)

# Atomic Terminology

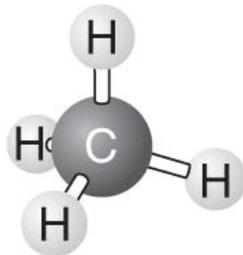
- **Molecules:** consist of two or more atoms ( $\text{H}_2\text{O}$ ,  $\text{CO}_2$ )



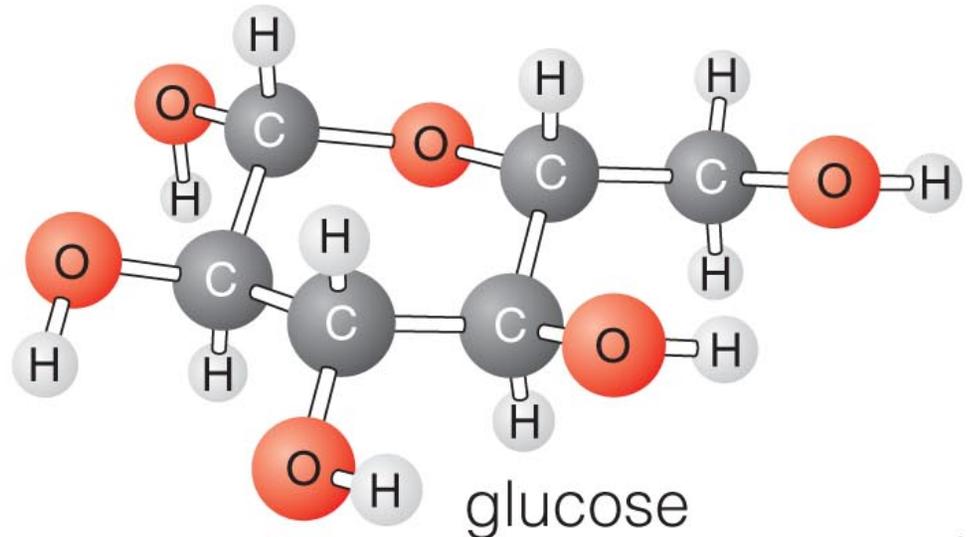
oxygen



water



methane



glucose

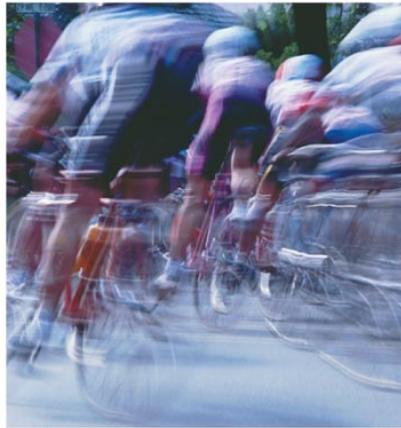
*Organic molecules contain carbon  
(and usually also contain hydrogen).*

*Compounds are molecules made from  
atoms of two or more different elements.*

*Molecules consist of two or more atoms.*

# What is energy?

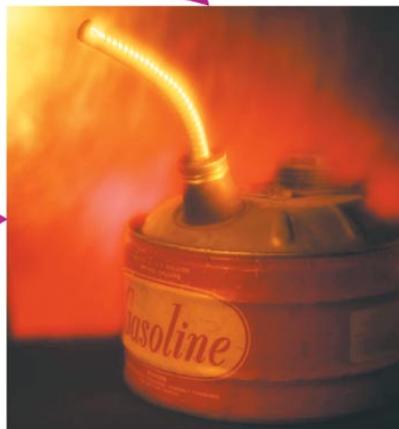
*Energy can be converted from one form to another.*



kinetic energy  
(energy of motion)



radiative energy  
(energy of light)



potential energy  
(stored energy)

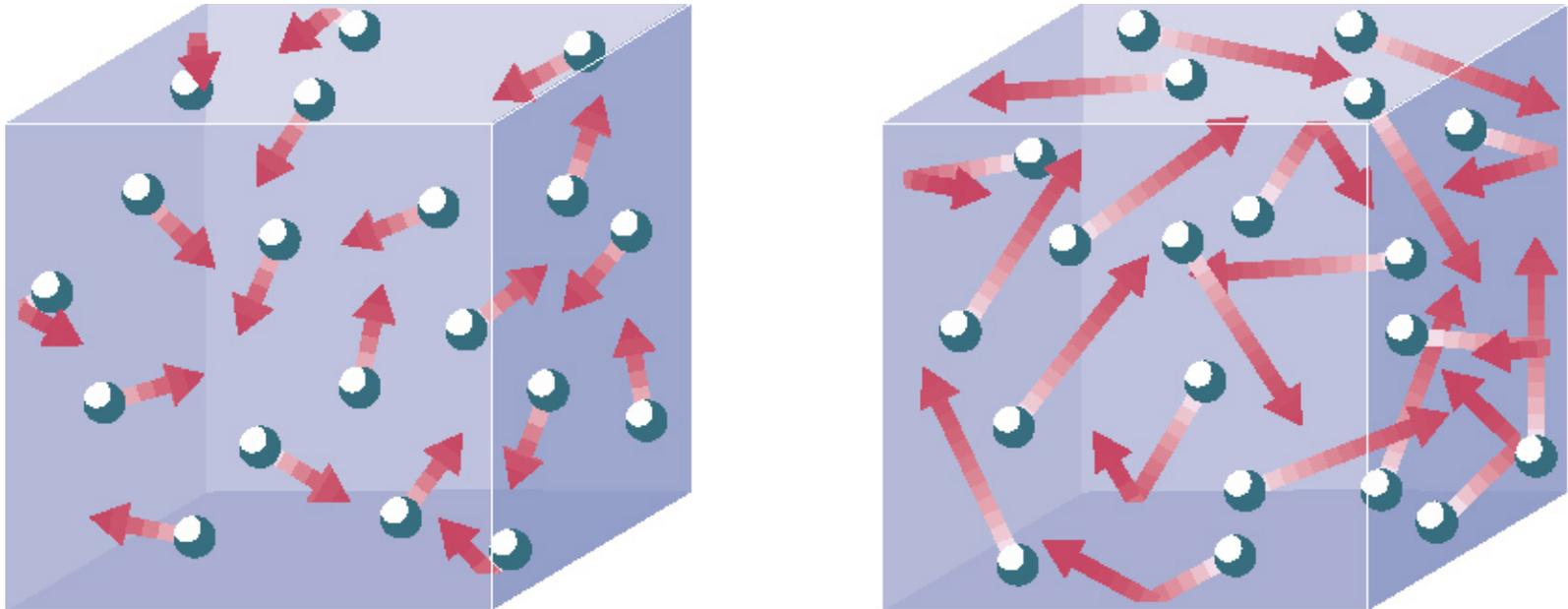
Energy is a very broad concept, basically it's anything that can make matter move or change

Energy changes forms but is not created or destroyed: this is a law of physics

Life mostly uses light energy (the Sun) and potential (chemical) energy from molecules

In all matter, anywhere in the universe, atoms and molecules in constant, microscopic motion

**Temperature** is a measure of the *average* kinetic energy of the many particles in a substance.



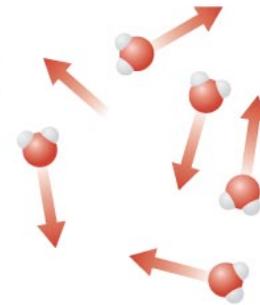
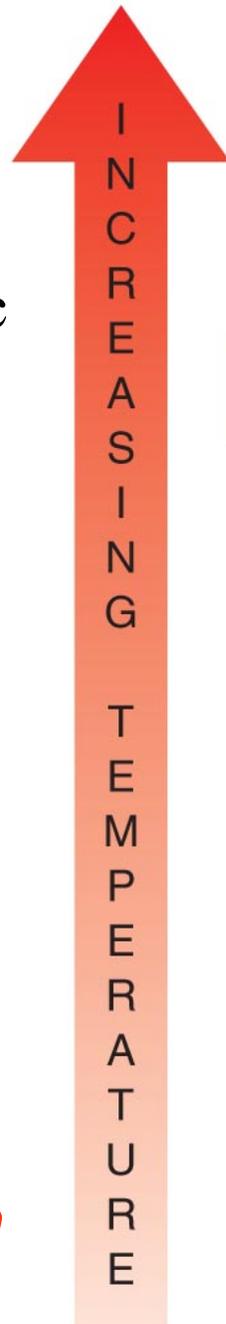
Longer arrows mean higher average speed.

All atoms and molecules in the universe are in constant microscopic motion or vibration:

### ***Thermal energy***

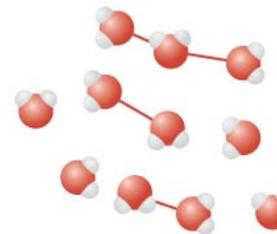
As a result, every substance emits a smooth spectrum of radiation, mostly at invisible infrared wavelengths:

### ***Thermal radiation***



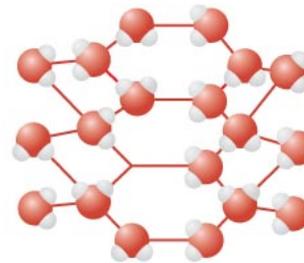
#### ***Gas Phase***

*Atoms or molecules move essentially unconstrained.*



#### ***Liquid Phase***

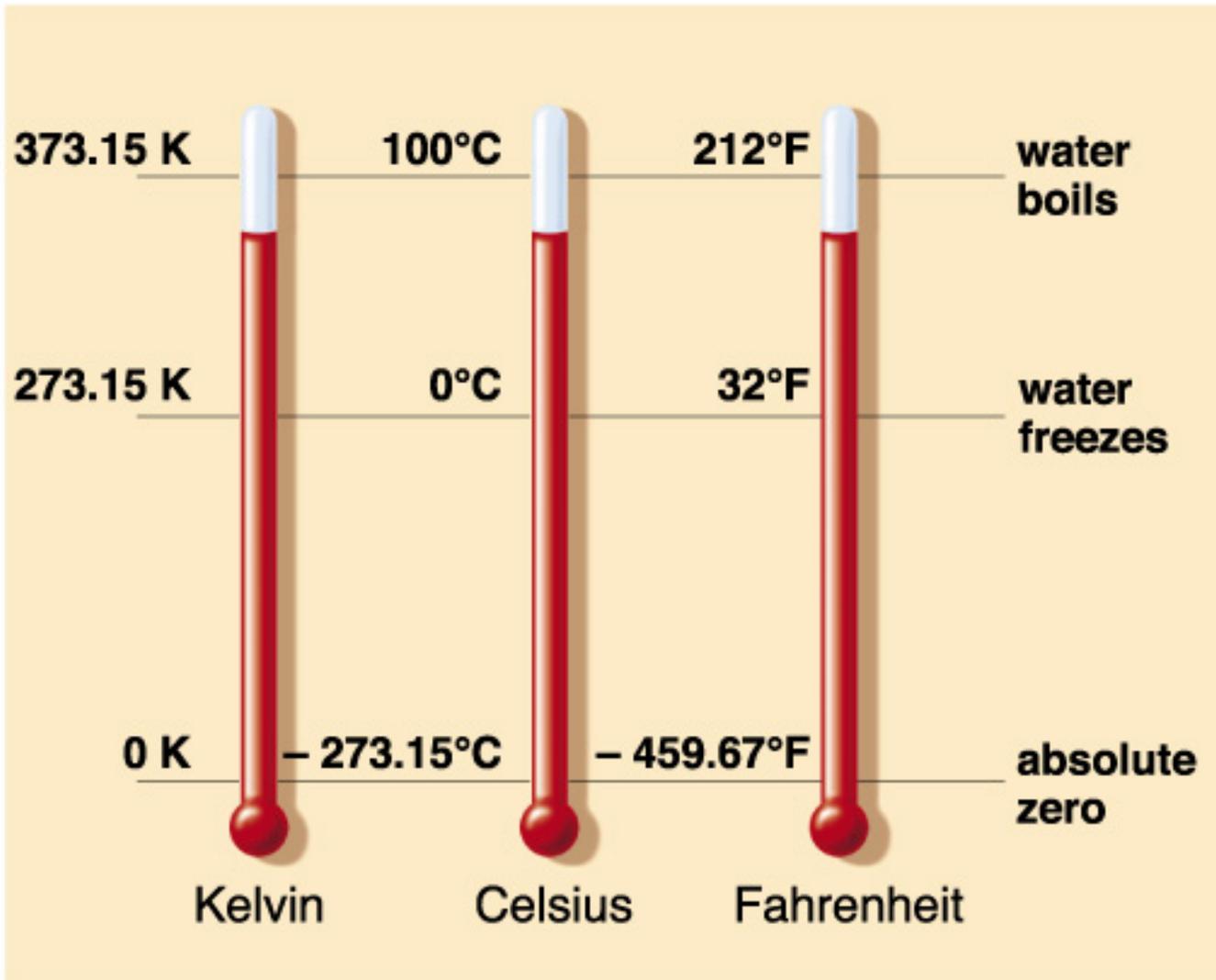
*Atoms or molecules remain together but move relatively freely.*



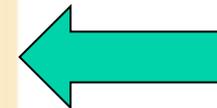
#### ***Solid Phase***

*Atoms or molecules are held tightly in place.*

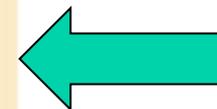
# Temperature Scales



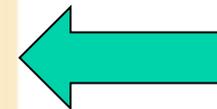
Stars are thousands of degrees



Terrestrial planets



Outer gas Planets



Most of the universe

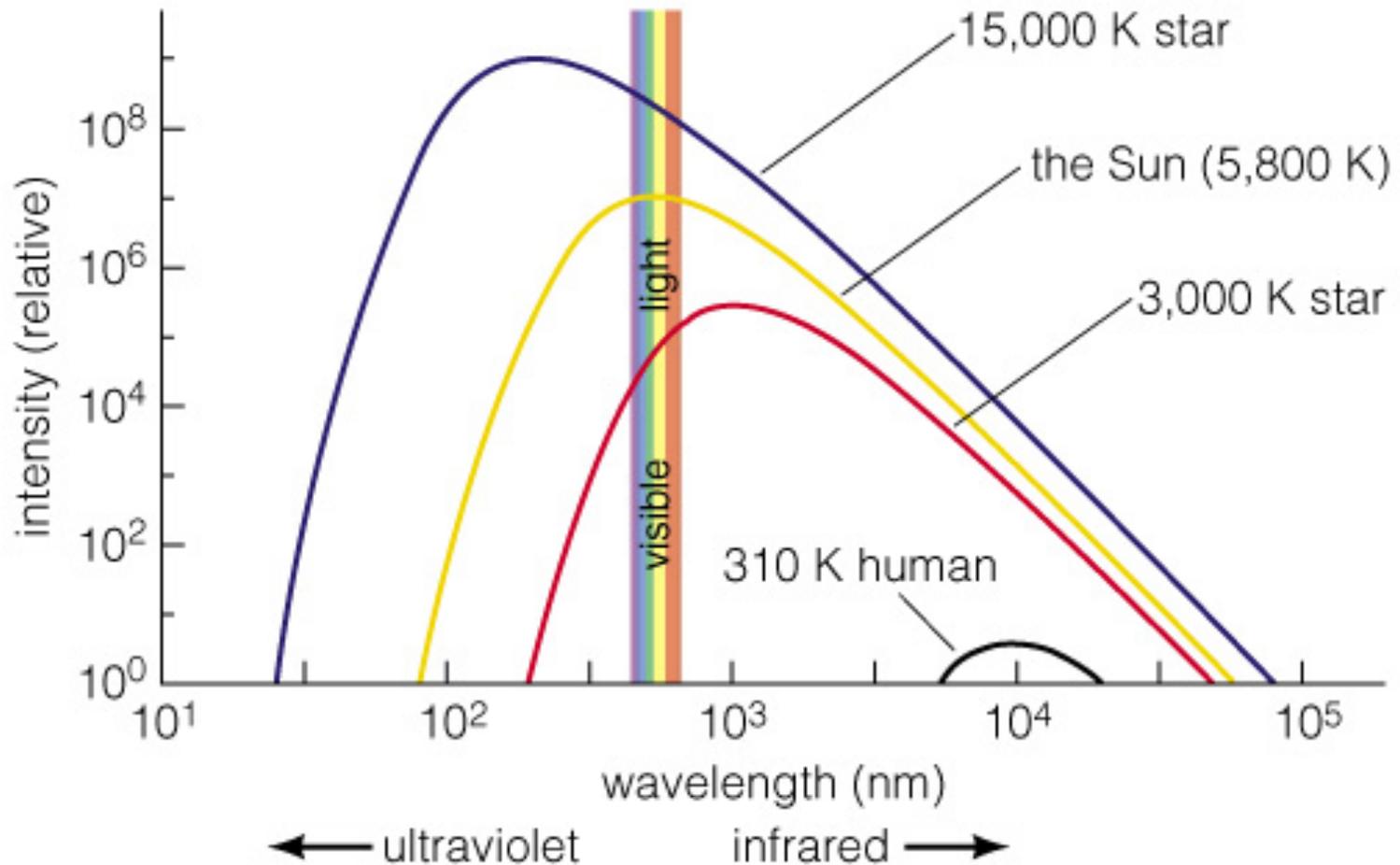
How does light tell us the temperatures of planets and stars?

## Thermal Radiation

- Nearly all large or dense objects emit thermal radiation, including stars, planets, you, me...
- Thermal radiation has a smooth and broad spectrum, like the Sun's rainbow of color
- An object's thermal radiation spectrum depends on only one property: its **temperature**

## Two Properties of Thermal Radiation:

1. Hotter objects emit more light at all frequencies per unit area.
2. Hotter objects emit photons with a higher average energy.



# How do light and matter interact?

- Emission
- Absorption
- Transmission
- Reflection or Scattering

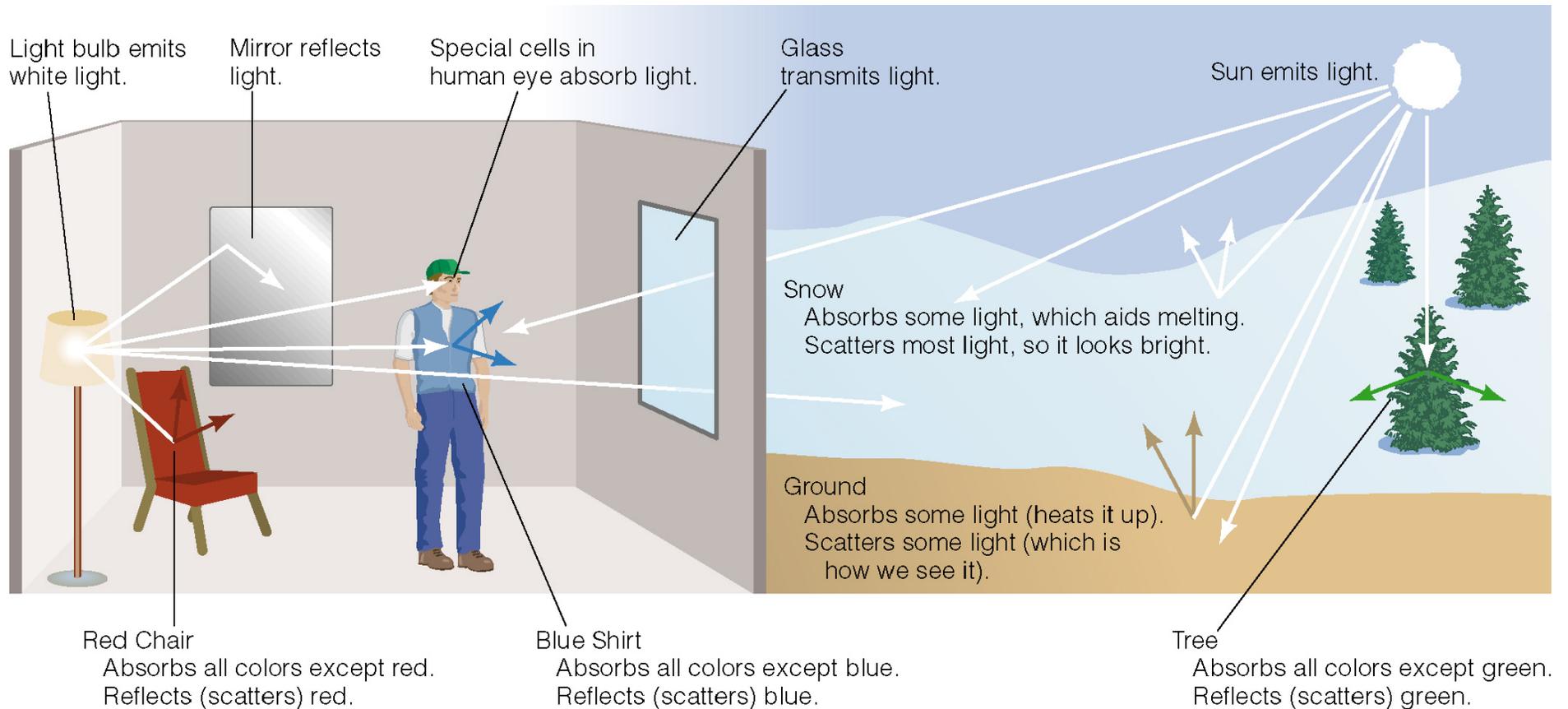
Everything we know  
about the universe is a  
result of these effects

## *Terminology:*

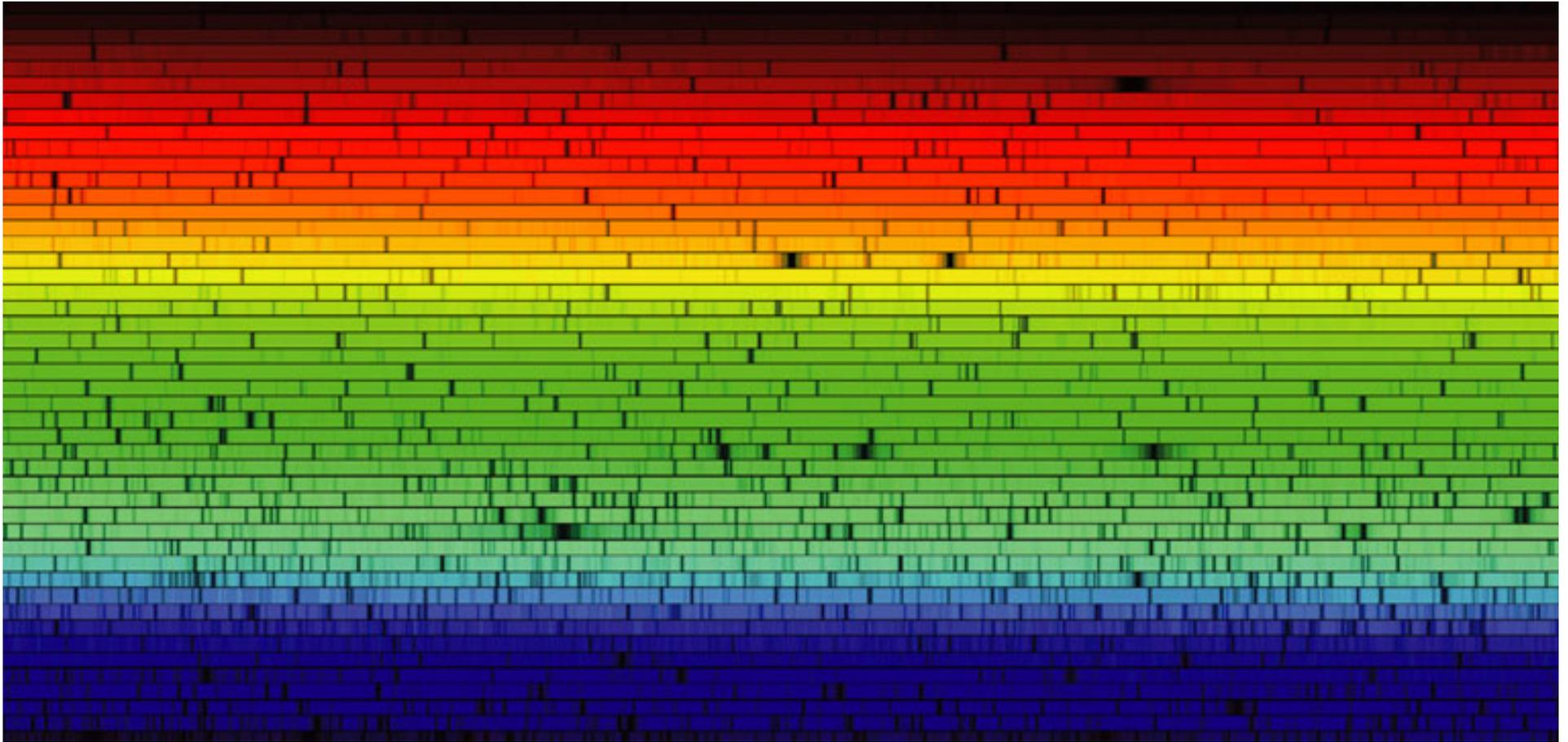
- Transparent: transmits light
- Opaque: blocks (absorbs) light

# Interactions of light and matter

Applied to a distant object like a planet, this can give us important physical information  remote sensing



# Example: the Sun's Spectrum

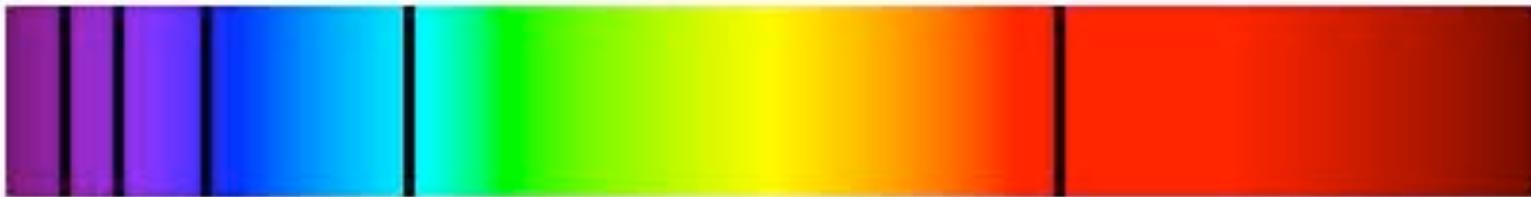


# Chemical Fingerprints

- Every atom, ion, and molecule has a unique spectral “fingerprint”
- We can identify the chemicals in gas by their fingerprints in the spectrum.
- With additional physics, we can figure out abundances of the chemicals, and often the temperature, pressure, and much more.

# Example: Hydrogen Spectrum

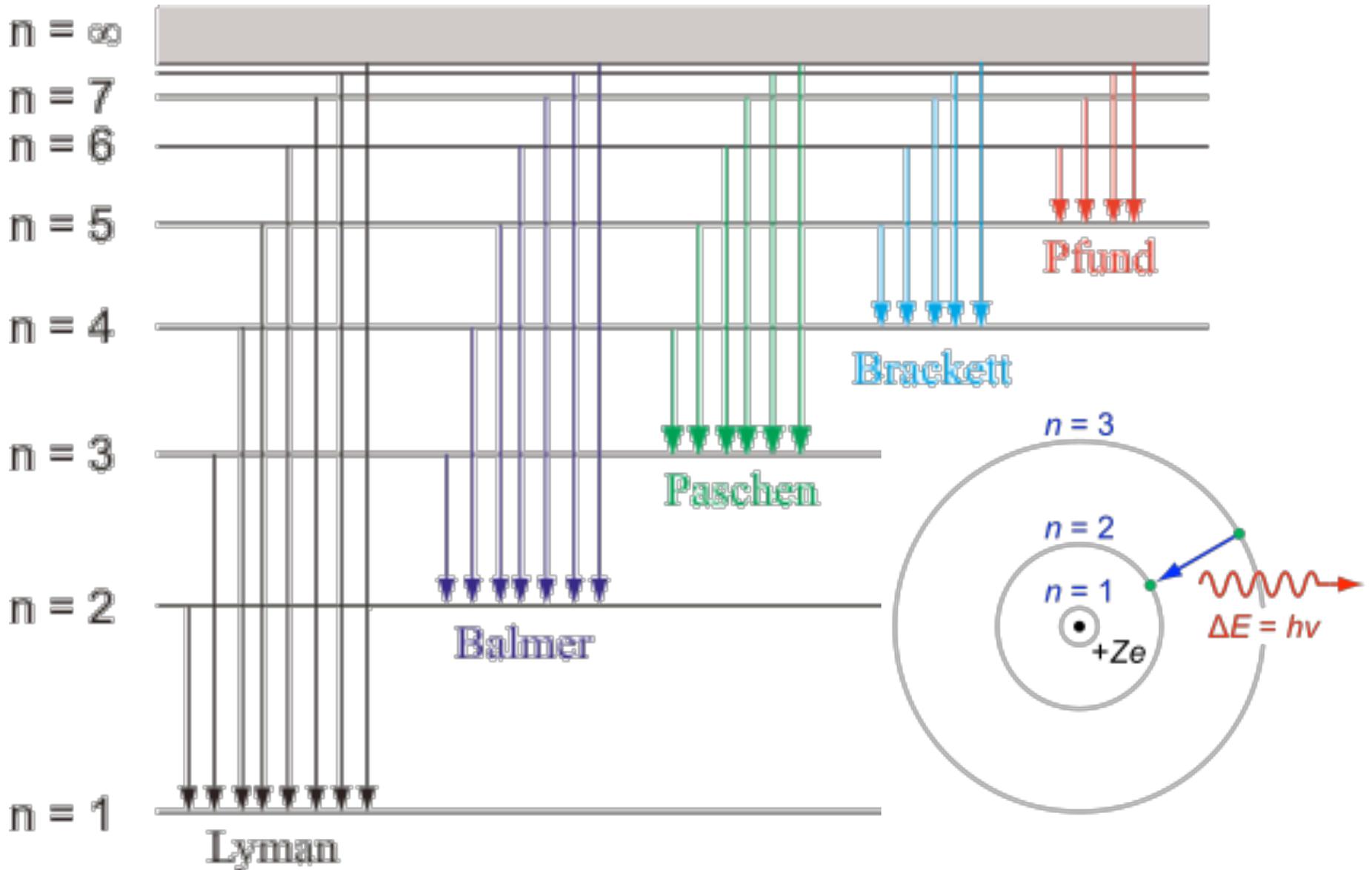
Hydrogen Absorption Spectrum



Hydrogen Emission Spectrum



# Hydrogen *Electronic* Energy Levels

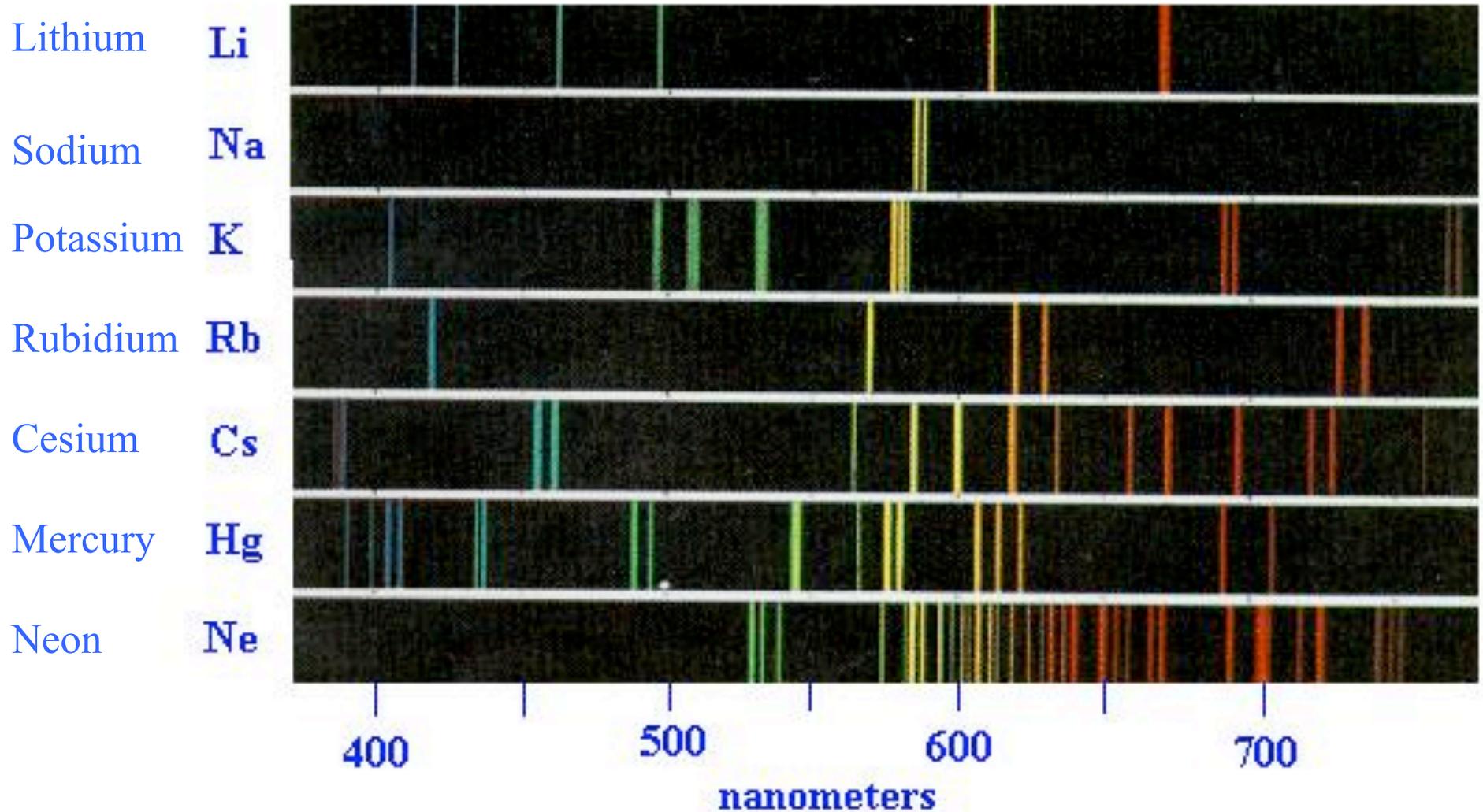


Hydrogen gas glowing (H $\alpha$  at 656 nm)



Credit: Nick Wright/IPHAS

# Example: More Elements...

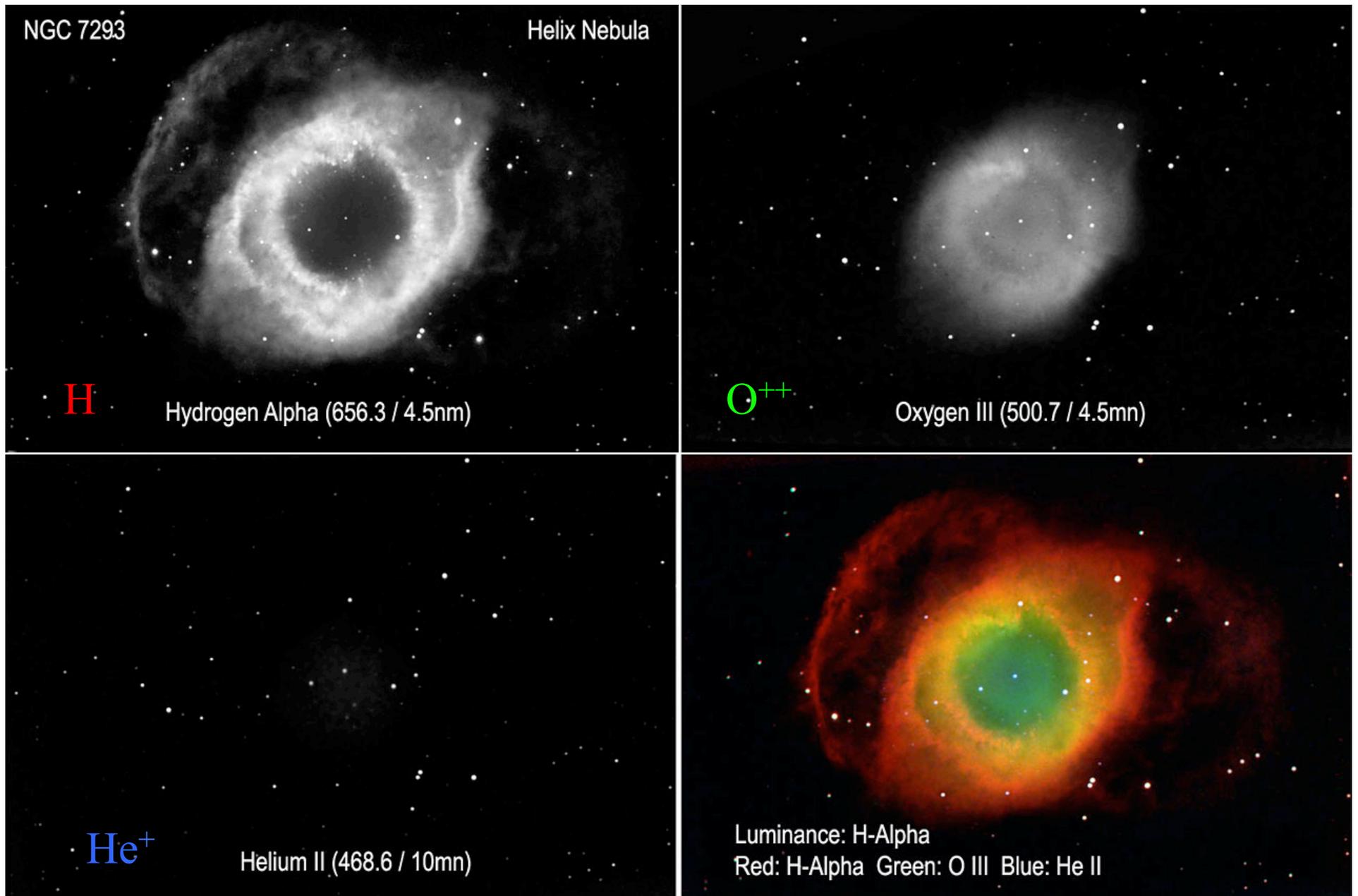


# Sodium gas glowing (yellow light)

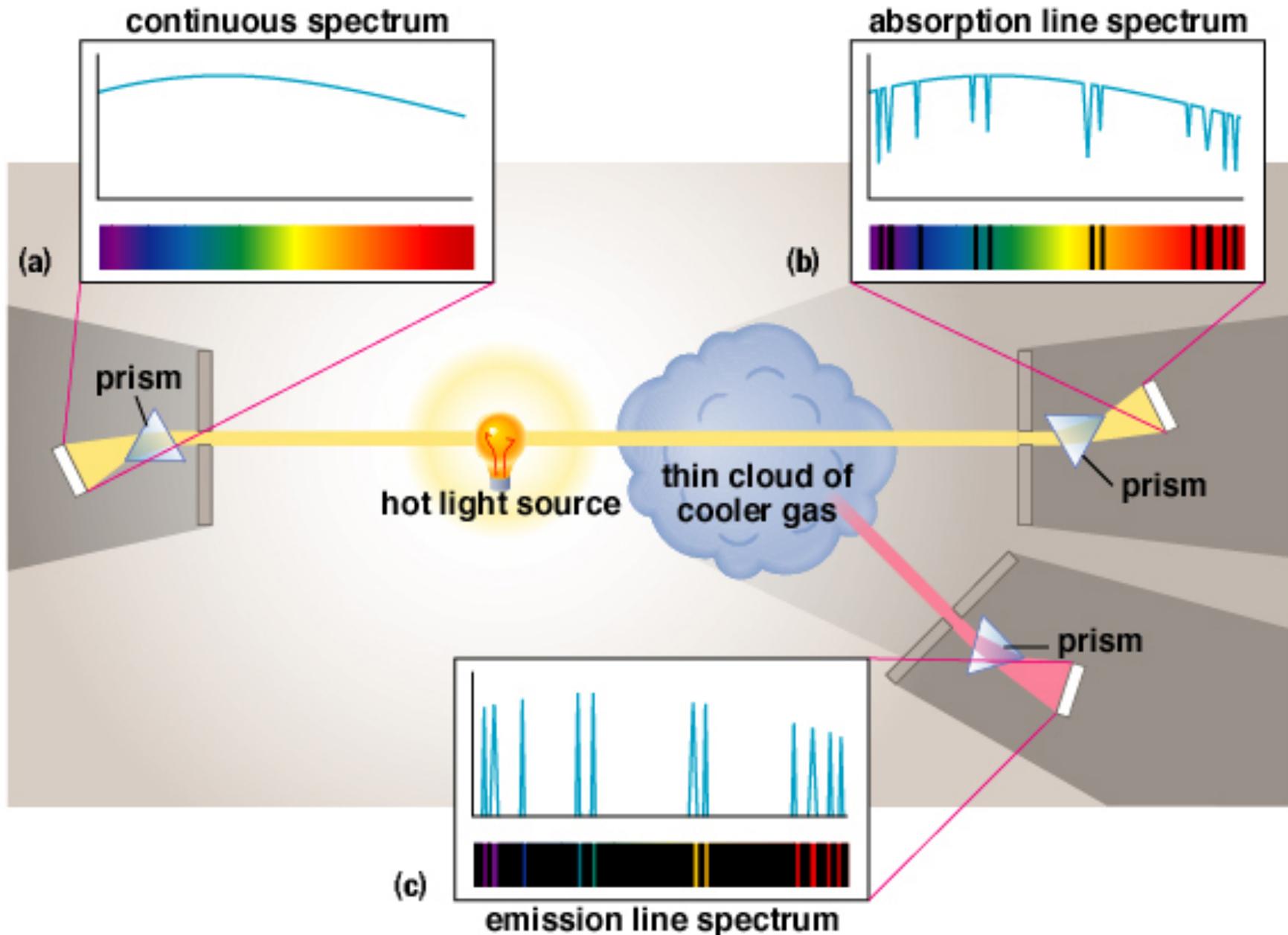


Look at previous slide to see why it is yellow...

# We can even see ionized elements!



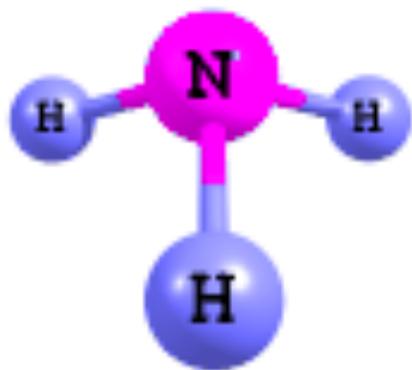
# What types of light spectra can we observe?



# We also see molecules in space

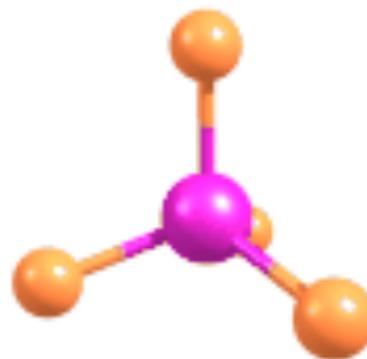
In addition to electronic transitions, molecules can also vibrate and rotate. These transitions are at lower energy and thus occur in the infrared to radio wavelengths.

Vibration



Vibrations are quantized

Rotation

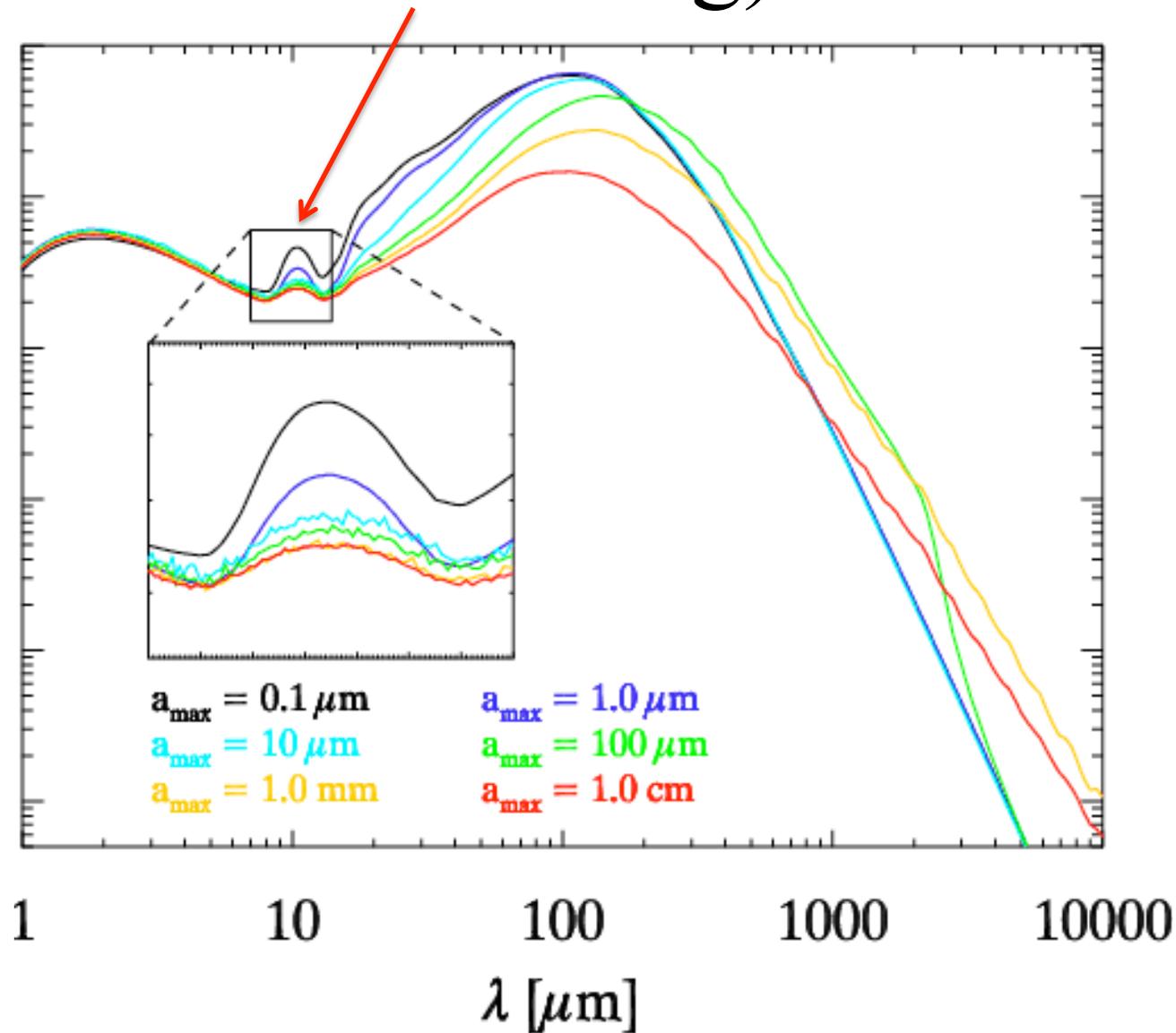


Rotations are also quantized

# Obscuring Dust Cloud ... what is it made of??



Dust is made primarily of Silicates  
(Si-O bonds vibrating)

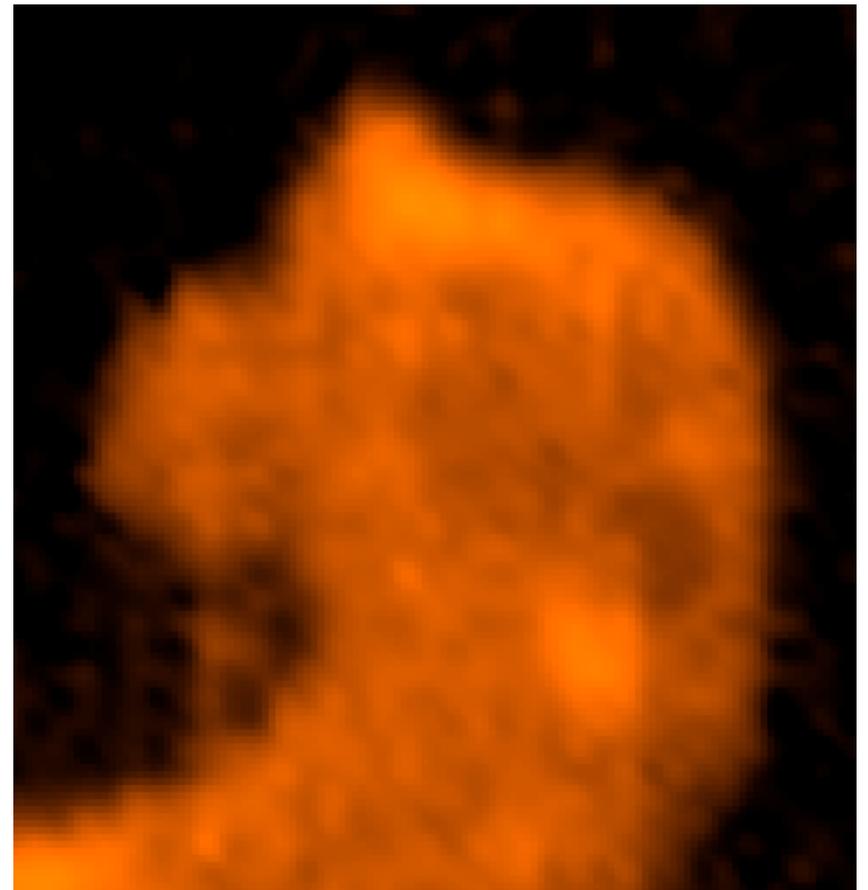


# Example: Rotational Transitions

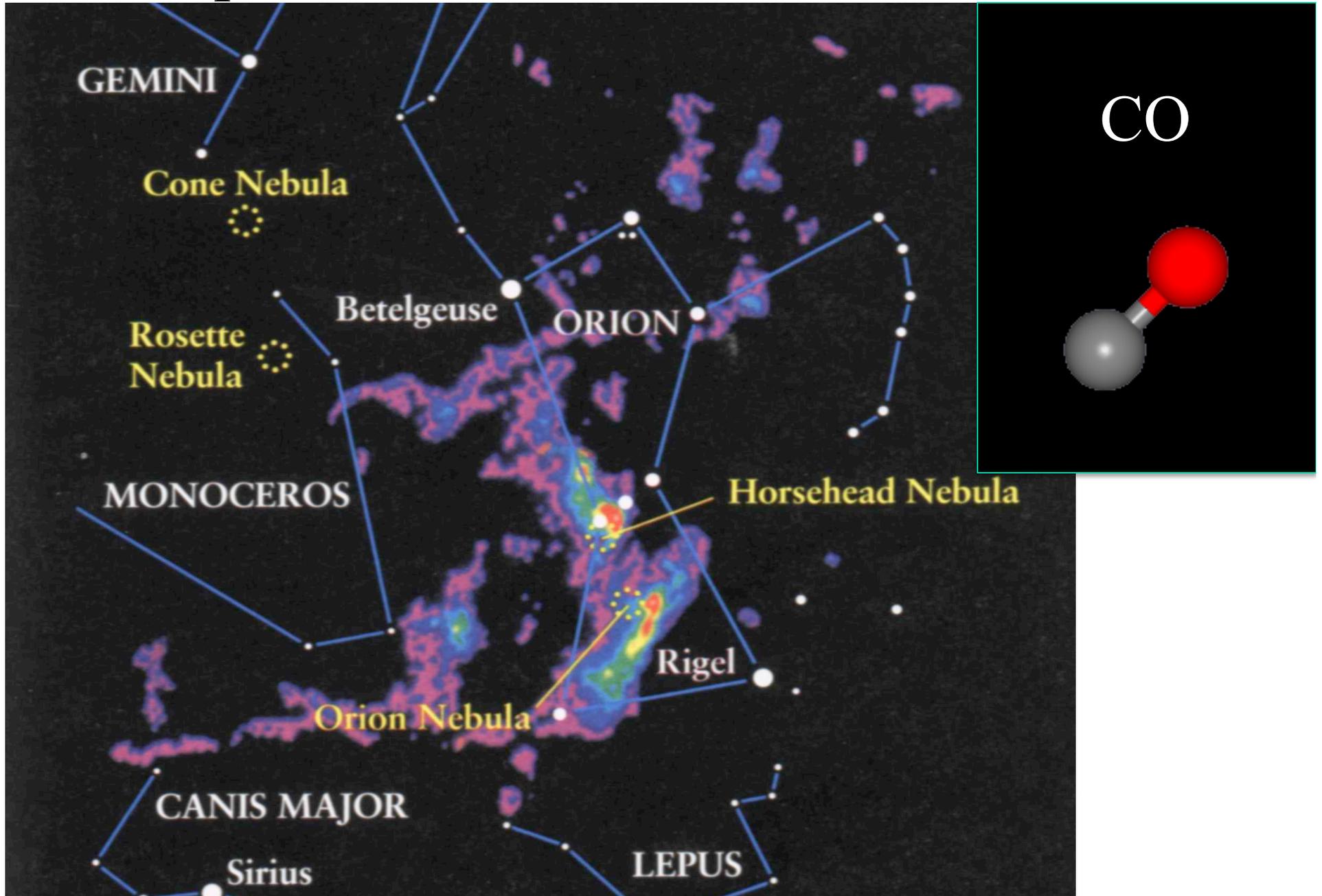
Optical Image



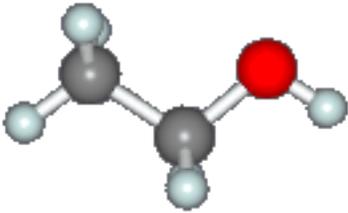
Carbon Monoxide  
molecules



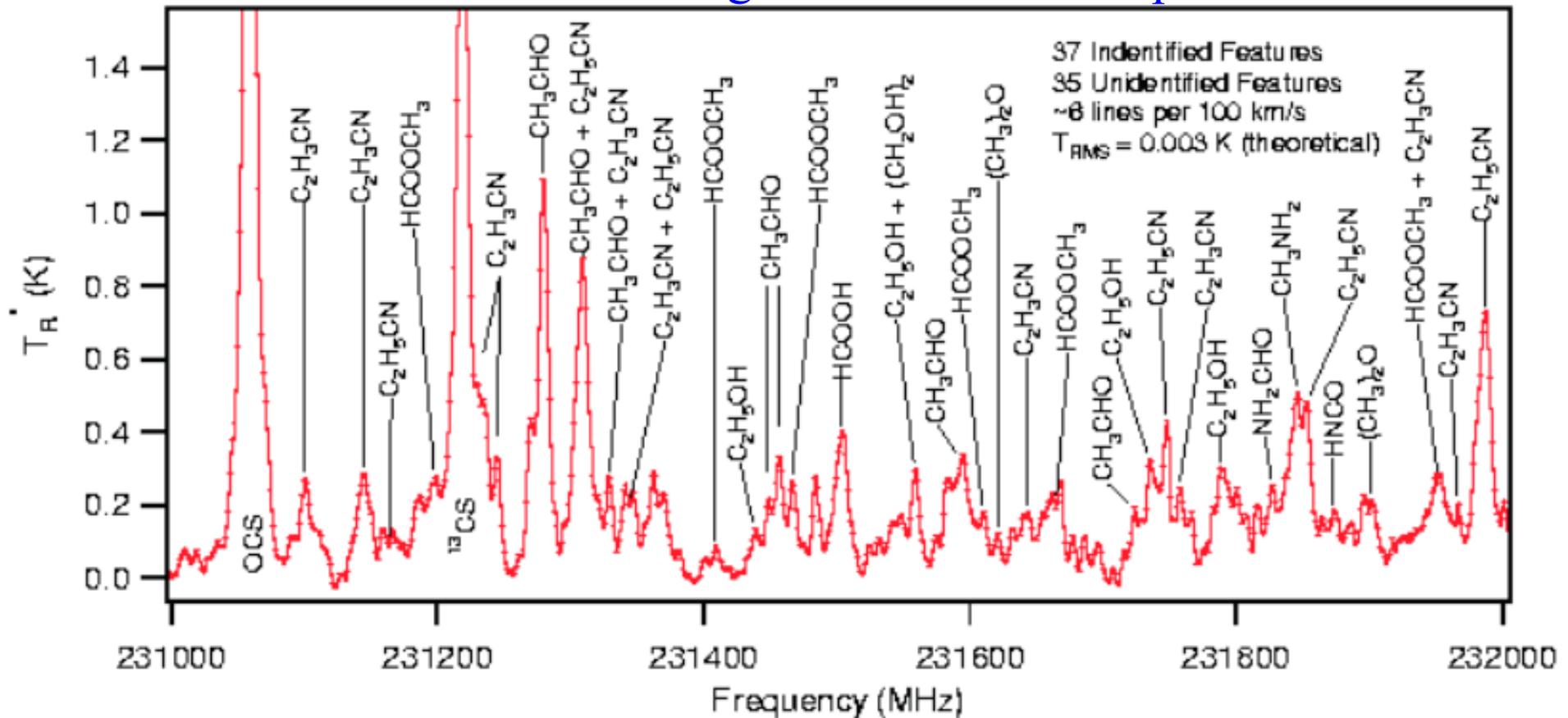
# Example: Carbon Monoxide towards Orion



# Rotational Transitions in Radio Spectrum

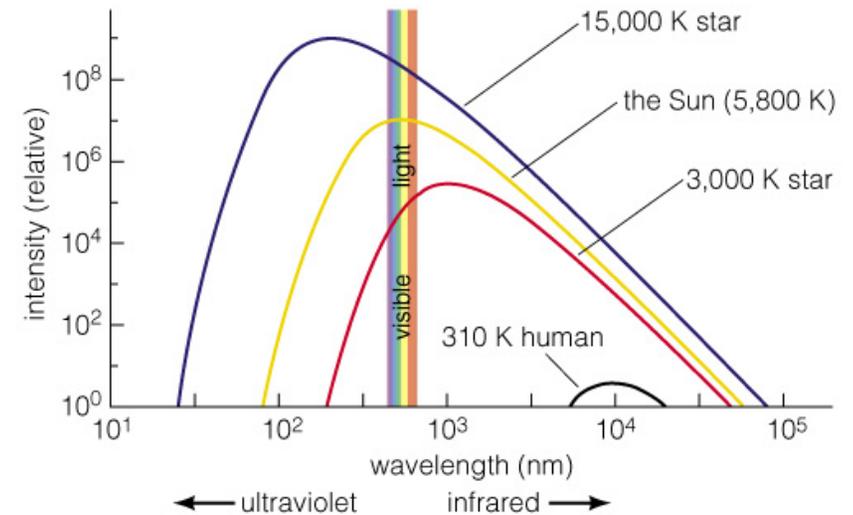
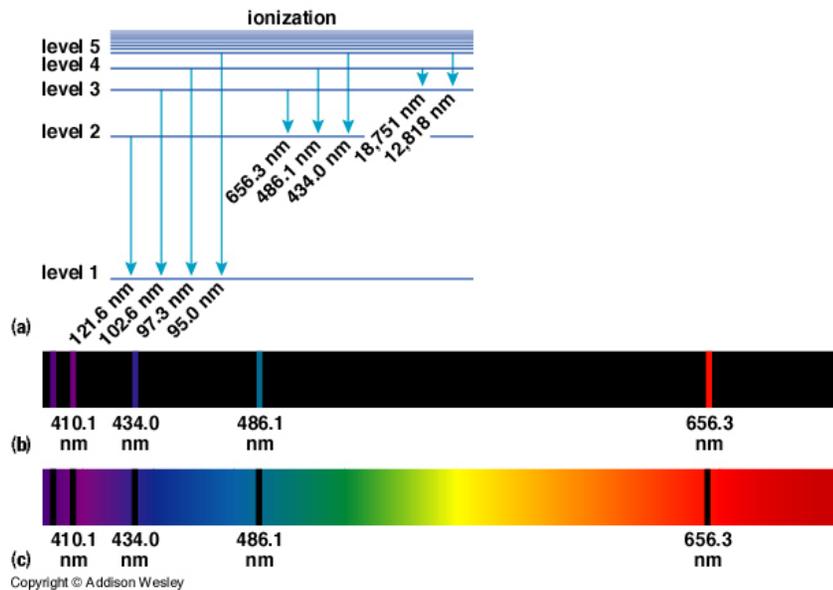


Detections of organic molecules in space!



Wavelength  $\sim 1\text{mm}$ . The frequency of this observations is  $\sim 2000\text{x}$  higher than FM radio

# Two Fundamentally Different Spectral Mechanisms



## Spectral lines:

- narrow, sharp features
- pattern unique to each element
- gives chemical composition
- emission (hot, diffuse gas)
- absorption (cool, diffuse gas)

## Thermal radiation:

- broad, smooth continuum
- peak emission gives temperature
- higher temp, shorter wavelength
- no information on composition
- everything emits thermally