

## AST 300B – Spring 2017

### In-class/take-home Problems Due: Wednesday April 12th

41. The emissivity coefficient  $j_\nu$  ( $\text{erg s}^{-1} \text{cm}^{-3} \text{ster}^{-1} \text{Hz}^{-1}$ ) for the two-photon transition  $2s - 1s$  in the Hydrogen atom has the same equation as for an emission line except that the line profile function is replaced by  $P_{2s}(\nu)$  which is the probability that a photon is emitted at frequency  $\nu$ . The Einstein  $A(2s-1s) \sim 8.23 \text{ s}^{-1}$ . A Hydrogen atom energy level diagram is on the back.

- a. In the low density limit, collisional excitation and de-excitation are negligible, and the only way to populate the  $2s$  level is by recombination at an effective recombination rate denoted by  $\alpha_{2s}^{\text{eff}}$ . Write down the expression for  $j_\nu$  in terms of  $P_{2s}(\nu)$  and  $\alpha_{2s}^{\text{eff}}$  assuming statistical equilibrium ( $dn_{2s}/dt = 0$ ). [Hints: You will need to use statistical equilibrium to relate  $n_{2s}$  to  $\alpha_{2s}^{\text{eff}}$ . Ignore Einstein B terms because we are in the high  $\nu$  limit.]
- b. It is possible that the density in HII regions can become high enough that the  $2s$  level is de-excited by collisions with protons and electrons to the  $2p$  level. The “downward” collision rates are given by  $\gamma(2s-2p) = 4.74 \times 10^{-4} \text{ cm}^3 \text{ s}^{-1}$  for collisions with protons and by  $\gamma(2s-2p) = 0.57 \times 10^{-4} \text{ cm}^3 \text{ s}^{-1}$  for collisions with electrons (Osterbrock 1974). What is the critical density ( $n_{\text{crit}} \text{ cm}^{-3}$ ) for the  $2s-1s$  transition?
- c. Given that the critical density in part (b) is not unreasonable for HII regions, we can't always assume that the low density limit applies. Modify your equation for  $j_\nu$  assuming that collisional de-excitation by protons and electrons from  $2s-2p$  cannot be ignored and assuming statistical equilibrium ( $dn_{2s}/dt = 0$ ). [Hint: You don't need to worry about collisions from  $2p-2s$  (because there is a **very fast** Einstein A from  $2p-1s$ ) or between  $2s$  and any other levels. Also, what can you assume about the number density of electrons and protons?]. Express your answer in terms of  $n_{\text{crit}}$ .

## Energy Levels of Hydrogen ( $n=1-4$ )

Shells K ( $n=1$ ,violet), L ( $n=2$ ,blue), M ( $n=3$ ,green) and N ( $n=4$ ,red)

