

Electrostatic Bremsstrahlung - "Free-Free" Radiation

ASTR
300B

In plasma, e^- and ion are accelerated as they pass each other on scales $<$ Debye length of plasma

The power radiated from an accelerated (\dot{v}) charge is given classically from Larmor's formula:

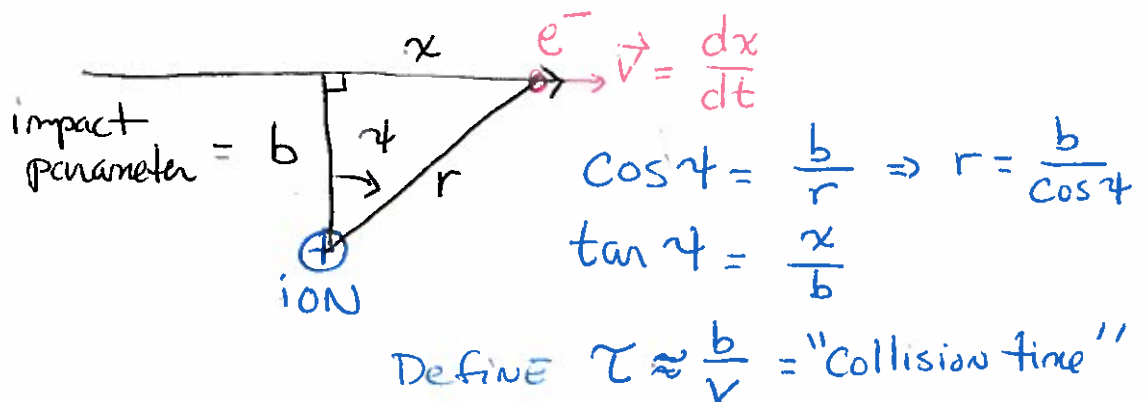
$$P = \frac{2}{3} \frac{e^2 \dot{v}^2}{c^3} \quad \text{erg} \cdot \text{s}^{-1} \quad \left\{ \begin{array}{l} \text{CGS NOTE:} \\ e \sim 4.8 \times 10^{-10} \text{ statcoulomb} \end{array} \right.$$

SINCE $F=ma$ then $a = \dot{v} \sim \frac{1}{m}$. $P \sim \dot{v}^2 \sim \frac{1}{m^2}$

This implies radiation from e^- much more important than radiation from ions because $P_e/P_{\text{ions}} \sim (m_p/m_e)^2 \sim 4 \times 10^6$.

Furthermore, interactions between identical particles are negligible because $\dot{v}_1 = -\dot{v}_2$ and the radiated electric fields are equal in magnitude but opposite in sign, so the net radiated field $\rightarrow 0$ at distances \gg collision impact parameter.

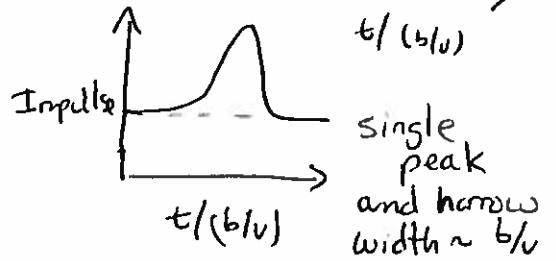
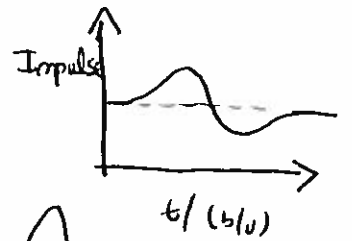
\Rightarrow ONLY e^- - ion collisions important
AND ONLY e^- radiate significantly



SINCE e^- moving so fast, we can approximate its trajectory as a straight line.

$$F_{||} = m_e \dot{V}_{||} = -\frac{ze^2}{r^2} \sin \psi = \frac{-ze^2 \sin \psi \cos^2 \psi}{b^2}$$

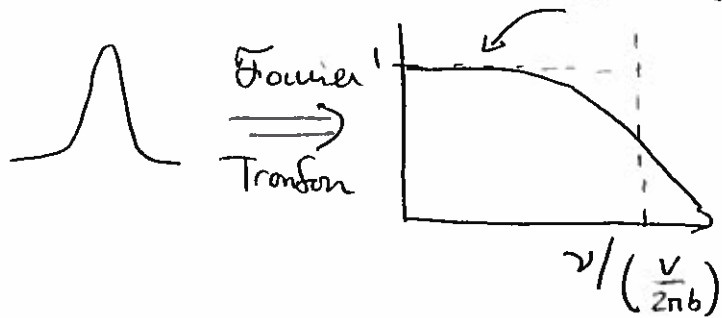
$$F_{\perp} = m_e \dot{V}_{\perp} = \frac{ze^2}{r^2} \cos \psi = \frac{ze^2 \cos^3 \psi}{b^2}$$



$\dot{V}_{||}$ pulse is sinusoidal with $\omega = 2\pi\nu \sim \frac{1}{\tau} = \frac{v}{b} \Rightarrow$ high ν (IR) radiation

\dot{V}_{\perp} pulse is single peaked (narrow width $\sim \frac{b}{v}$)

Fourier transform is \approx flat up to a $\nu \sim \frac{v}{2\pi b}$



\Rightarrow Frequency response is nearly const with ν at low frequency.

For rest of problem, we will only consider \dot{V}_{\perp}

Our derivation will get correct functional dependence, but will not get constants correct. To do this problem correctly need to:

① Calculate electric field $E(t)$ of far field electric dipole

② Take Fourier Transform $|\tilde{E}_{\nu}|$

③ Integrate $|\tilde{E}_{\nu}|^2$ over solid angle to get W_{ν} erg $\cdot \text{Hz}^{-1}$

Solution is: $W_{\nu} = \frac{4\pi^2}{3} \frac{ze^6}{m_e^2 c^3 b^2 v^2} e^{-4\pi\nu b/v}$

Let's work through the simpler derivation ...

Power radiated from accelerated \dot{v}_\perp change :

$$P_\perp = \frac{2}{3} \frac{e^2 \dot{v}_\perp^2}{c^3} = \frac{2e^2}{3c^3} \frac{z^2 e^4}{m_e^2} \left(\frac{\cos^3 \gamma}{b^2} \right) \text{ erg} \cdot \text{s}^{-1}$$

The total energy emitted by the pulse is :

$$W = \int_{-\infty}^{\infty} P_\perp dt \text{ erg}$$

Strategy: use geometry and approximately const $\vec{e} \cdot \vec{v}$ to turn $dt \rightarrow d\gamma$

$$v = \frac{dx}{dt} = b \frac{d(\tan \gamma)}{dt} = b \frac{\sec^2 \gamma d\gamma}{dt} \Rightarrow dt = \frac{b}{v} \frac{d\gamma}{\cos^2 \gamma}$$

$$W \sim \frac{z^2 e^6}{c^3 m_e^2 b^3 v} \int_0^{\pi/2} \cos^4 \gamma d\gamma \quad \xrightarrow{3\pi/16} \text{just a constant}$$

Now, since we know that Fourier transform of pulse profile was nearly flat b/w to $\nu_{\max} = \frac{1}{2\pi\tau} = \frac{v}{2\pi b}$

Let's approximate W_ν (average energy emitted per unit frequency)

$$W_\nu \approx \frac{W}{\nu_{\max}} \approx \frac{W}{\left(\frac{v}{2\pi b}\right)} \sim \frac{z^2 e^6}{c^3 m_e^2 b^3 v} \cdot \frac{b}{v}$$

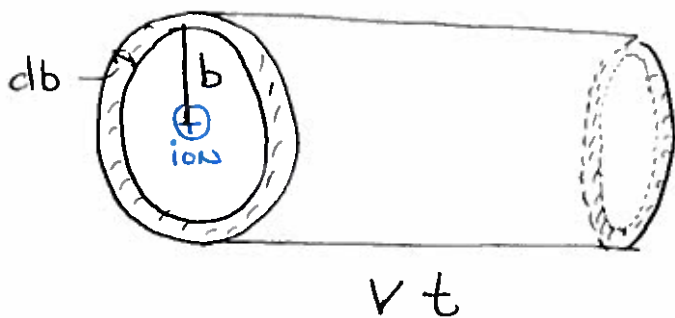
$$W_\nu \sim \frac{z^2 e^6}{c^3 m_e^2 b^2 v^2} \quad \text{erg} \cdot \text{Hz}^{-1} \quad \text{NOTE: } W_\nu(b, v)$$

Valid for $\nu < \nu_{\max} \approx 10^{14}$ Hz in HII regions.

This is for a single interaction \Rightarrow

Need to $\iint_{b, v}$ to get $j_\nu \dots$

In LTE $(KE)_{ions} = (KE)_{e^-} \Rightarrow$ ions are essentially stationary during e^- interaction



The number of e^- passing a single ion per unit time at impact parameters $b + db$ with speed $v + dv$ is

THINK: $\frac{1}{t} = n e v$

$$n_e (2\pi b db) v f(v) dv \quad s^{-1}$$

\uparrow speed distribution of e^-
IN LTE = Maxwell-Boltzmann Distribution $\int f(v) dv = 1$

This is only for 1 target ion. To get total number of "collisions" per unit volume multiply by $n_i \text{ cm}^{-3}$

The free-free emissivity (spectral power per 4π ster):

$$j_\nu = \frac{1}{4\pi} \int_{b=0}^{\infty} \int_{v=0}^{\infty} W_\nu(v, b) n_i n_e 2\pi b db v f(v) dv$$

$$\sim \frac{Z^2 e^6 n_e n_i}{c^3 m_e^2} \int_{v=0}^{\infty} \frac{f(v)}{v} dv \int_{b=0}^{\infty} \frac{db}{b}$$

Let's evaluate these 2 integrals:

First, the velocity integral :

$$f(v) = 4\pi \left(\frac{m_e}{2\pi kT} \right)^{3/2} e^{-m_e v^2 / 2kT} \cdot v^2$$

We have to consider the additional issue of discreteness of photon energies: $h\nu = \frac{1}{2} m_e v_{\min}^2$ sets v_{\min}

In other words, e^- must have at least v_{\min} to produce photons of energy $\nu \rightarrow$ thus set the lower limit to \int .

$$\int_{v_{\min}}^{\infty} \frac{f(v)}{v} dv \sim \left(\frac{2m_e}{\pi kT} \right)^{1/2} e^{-m_e v_{\min}^2 / 2kT}$$

$$\sim \left(\frac{2m_e}{\pi kT} \right)^{1/2} e^{-h\nu / kT}$$

At radio ν , $e^{-h\nu / kT} \approx 1$ for $T = 10^4$ K H II regions.
This term only important at high ν .

Notice the integral over db diverges logarithmically!

Thus means we have to choose physically meaningful limits b_{\min} & b_{\max}

$$\int_{b_{\min}}^{b_{\max}} \frac{db}{b} = \ln \left(\frac{b_{\max}}{b_{\min}} \right)$$

b_{\min}

b_{\max}

Classical: Maximum momentum $m_e \Delta v$ is $2 \times$ initial momentum

$$m_e \Delta v = \int_{-\infty}^{\infty} F_{\perp} dt = \frac{2Ze^2}{bv}$$

$$\Rightarrow b_{\min} \sim Ze^2 / m_e v^2$$

Quantum: Heisenberg Uncertainty

$$b_{\min} \approx \hbar / m_e v$$

Classical: Largest b that can generate significant power at ν :

$$b_{\max} \sim \frac{v}{\omega} = \frac{v}{2\pi\nu}$$

See Fourier Transform Figure

Debye length $\lambda_D \approx \left(\frac{kT}{4\pi n_e e^2} \right)^{1/2}$

Characteristic length scale for charge shielding in a plasma.

For H II regions, classical limits are appropriate:

$$\frac{b_{\max}}{b_{\min}} \sim \frac{v}{2\pi v} \cdot \frac{m_e v^2}{Z e^2} \sim \frac{1}{v}$$

$$\Rightarrow \int_{b_{\min}}^{b_{\max}} \frac{db}{b} \sim \ln\left(\frac{1}{v}\right) \sim v^{-0.1} \text{ is reasonable approximation.}$$

The more precise calculation takes quantum effects into account and averages over $f(v)$. It is called the Gaunt factor

$$\overline{g_{ff}} \approx \frac{\sqrt{3}}{\pi} \left[\ln \frac{(2kT)^{3/2}}{\pi Z e^2 m_e^{1/2} v} - \frac{5\gamma}{2} \right]$$

Euler's Constant

$$\approx 6.155 (Z v_{\text{GHz}})^{-0.118} (T/10^4 \text{K})^{0.177}$$

see Draine Chapter 10

So the emissivity coefficient is then: Using results from "correct" observation now

$$j_\nu = \frac{8}{3} \left(\frac{2\pi}{3}\right)^{1/2} \frac{Z^2 e^6}{m_e^2 c^3} \left(\frac{m_e}{kT}\right)^{1/2} n_e n_i e^{-h\nu/kT} \overline{g_{ff}}$$

erg. s⁻¹. cm³. ster⁻¹. Hz⁻¹

Notice at low ν (when $e^{-h\nu/kT} \sim 1$) that

$$j_\nu \sim \nu^{-0.118} \leftarrow \text{very weak frequency dependence.}$$

This is because as e^- flies by ion, its acceleration is like a sharp pulse. The Fourier transform of a sharp pulse is \approx flat with frequency.