

AST 250 – Spring 2018

Homework Due: Wednesday April 25

41. The CMB originates when the Universe was ~ 3000 K. Today, the CMB is observed to be 2.725 K. What redshift corresponds to the “surface of last scattering” of the CMB?
42. The density parameter, Ω_m , measures the density of matter in the Universe relative to the “critical density”, ρ_{crit} , defined as the density for which a mass shell in the Universe is expanding at its escape velocity. Let’s first calculate the critical density by setting the velocity of the Hubble expansion, $v = H d$, equal to the escape speed for a spherical volume of space with radius d and total enclosed mass M . Convert mass to density and solve for the density – this is defined as the critical density, ρ_{crit} . Ω_m is then defined as the ratio of the observed mass density, ρ_0 , to the critical density: $\Omega_m = \rho_0/\rho_{\text{crit}}$. Now write down the final expression for Ω_m in terms of ρ_0 , G , H , and numerical constants. The latest results from Planck CMB measurements indicate that $\Omega_m = 0.3089$, of which 15% is normal baryonic matter. Calculate the density of H atoms in the visible Universe (answer in units of number H atoms per m^3 . Assume all baryonic matter is Hydrogen).

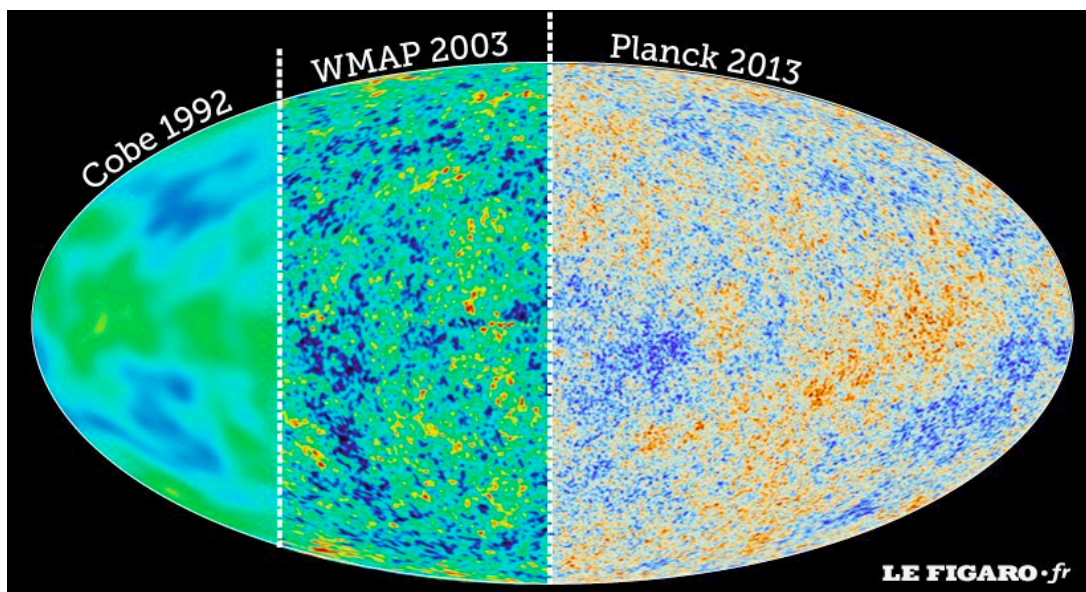


Figure 1: 3 of the space missions to measure anisotropy in the CMB.