

Albedo

ASTR 250

Planets and small bodies such as asteroids, comets, etc. reflect light at optical λ . They are not perfect blackbody absorbers. The fraction of light reflected is called the Albedo (denoted A)

Not to be confused with extinction A_λ !

The flux incident on a planet is:

$$F_{\text{incident}} = \frac{L_*}{4\pi d_*^2} \leftarrow \text{distance between planet and star}$$

$\text{erg} \cdot \text{s}^{-1} \cdot \text{cm}^{-2}$

$$\left(\frac{dE}{dt}\right)_{\text{reflect}} \equiv L_{\text{reflect}} = A \cdot \pi R_{\text{pl}}^2 \cdot \frac{L_*}{4\pi d_*^2}$$

$\text{erg} \cdot \text{s}^{-1}$

Albedo \times $\underbrace{\quad}_{\text{cross-section area}}$ \times incident flux

$$\left(\frac{dE}{dt}\right)_{\text{absorb}} = (1 - A) \cdot \pi R_{\text{pl}}^2 \cdot \frac{L_*}{4\pi d_*^2}$$

\uparrow
reflect + absorb = 100%

②

Let's assume that the planet re-radiates the absorbed energy like a blackbody at T_{PI} :

$$L_{planet} = 4\pi R_{PI}^2 \sigma T_{PI}^4$$

In equilibrium, $L_{planet} = \left(\frac{dE}{dt}\right)_{absorb}$ and we can calculate T_{PI} :

$$4\pi R_{PI}^2 \sigma T_{PI}^4 = (1-A) \pi R_{PI}^2 \frac{L_*}{4\pi d_*^2}$$

$$T_{PI}^4 = \frac{1-A}{4\sigma} \cdot \frac{4\pi R_*^2 \sigma T_*^4}{4\pi d_*^2}$$

$$T_{PI} = T_* \cdot \left(\frac{R_*}{d_*}\right)^{1/2} \cdot \left(\frac{1-A}{4}\right)^{1/4}$$

Note - assumes thermal energy is distributed evenly over $4\pi R_{PI}^2$ surface area (fast rotating or thick atmosphere...)

Plugging in values for Earth/Sun: (using $A \sim 0.37$)

$$T_{\oplus} \sim 248 \text{ K} \sim -13^\circ \text{F} \quad \text{Below Freezing!}$$

The Earth is actually warmer because molecules in the atmosphere (H_2O , CO_2 , CH_4 , etc.) can absorb IR radiation and effectively act as a blanket = Greenhouse Effect.

Without some greenhouse warming, we would be an ice ball!

The Habitable Zone is calculated by

solving for d_* when $T_{pl} = 273 - 373 \text{ K}$

(freezing - boiling of H_2O @ STP)

In reality, you need atmospheric models to account for greenhouse effect on T_{pl} .

$$r_{HZ} \sim \frac{\sqrt{(1-A) L_*}}{T_{pl}^2} \sim L_*^{1/2}$$