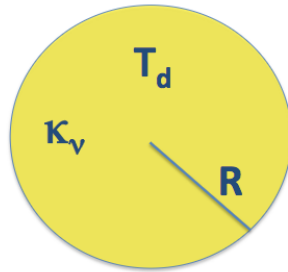


## AST 300B – Spring 2019

### In-class/take-home Problems Due: Monday Feb. 11

15. (a) Consider a spherical, optically thin dust thermal cloud with constant dust temperature  $T_d$ , constant dust opacity  $\kappa_v$ , and a radius  $R$ . Using the results from a previous homework, derive an equation relating the observed flux density of this cloud at a distance  $d \gg R$  to the mass of the cloud  $M$ . Write your answer in terms of the mass opacity  $\kappa_v$  of the cloud.



(b) Consider a millimeter mapping survey of the Rho Ophiuchus molecular cloud at  $\lambda = 1.3$  mm (below). If the cloud is located at a distance of 135 pc (the nearest molecular cloud to the Earth) and the mapping has a 3-sigma flux density sensitivity of 0.1 Jy (N.B.  $1 \text{ Jy} = 10^{-23} \text{ erg s}^{-1} \text{ cm}^{-2} \text{ Hz}^{-1}$ ), what **gas mass** is the survey sensitive to (quote your answer in solar masses) for dense cores at  $T_d = 10\text{K}$  and  $20 \text{ K}$ . If dense cores are up to 50% efficient (meaning  $\leq 50\%$  of their mass goes into a forming star), could this survey detect dense cores capable of forming stars down to the hydrogen burning limit ( $M \sim 0.08 M_{\text{sun}}$ )?  $\kappa_v \sim 0.9 \text{ cm}^2/\text{g}$  of dust at 1.3mm (OH5).

